

Data Set Catalog #143

OGO-5, Spectrometer

68-014A-18A

14 tapes

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1. INTRODUCTION:

The documentation for this data set was originally on paper, kept in NSSDC's Data Set Catalogs (DSCs). The paper documentation in the Data Set Catalogs have been made into digital images, and then collected into a single PDF file for each Data Set Catalog. The inventory information in these DSCs is current as of July 1, 2004. This inventory information is now no longer maintained in the DSCs, but is now managed in the inventory part of the NSSDC information system. The information existing in the DSCs is now not needed for locating the data files, but we did not remove that inventory information.

The offline tape datasets have now been migrated from the original magnetic tape to Archival Information Packages (AIP's).

A prior restoration may have been done on data sets, if a requestor of this data set has questions; they should send an inquiry to the request office to see if additional information exists.

2. ERRATA/CHANGE LOG:

NOTE: Changes are made in a text box, and will show up that way when displayed on screen with a PDF reader.

When printing, special settings may be required to make the text box appear on the printed output.

Version	Date	Person	Page	Description of Change
01				
02				

3 LINKS TO RELEVANT INFORMATION IN THE ONLINE NSSDC
INFORMATION SYSTEM:

<http://nssdc.gsfc.nasa.gov/nmc/>

[NOTE: This link will take you to the main page of the NSSDC Master Catalog. There you will be able to perform searches to find additional information]

4. CATALOG MATERIALS:

- a. Associated Documents To find associated documents you will need to know the document ID number and then click here.
<http://nssdcftp.gsfc.nasa.gov/miscellaneous/documents/>

- b. Core Catalog Materials

OGO 5

O, HE + H ION CONCS + EPHEM, TAPE

68-014A-18A

THIS DATA SET HAS BEEN RESTORED. THERE WERE ORIGINALLY FOURTEEN 7-TRACK, 800 BPI TAPES, WRITTEN IN BINARY. THERE ARE TWO RESTORED TAPES, WHICH WERE PADDED DURING THE RESTORATION PROCESS. THE DR TAPES ARE 3480 CARTRIDGES AND THE DS TAPES ARE 9-TRACK, 6250 BPI. THE ORIGINAL TAPES WERE CREATED ON A 1108 COMPUTER. THE DR AND DS NUMBERS ALONG WITH THE CORRESPONDING D NUMBERS AND TIME SPANS ARE AS FOLLOWS:

DR#	DS#	D#	FILES	TIME SPAN
DR003945	DS003945	D006071	1-10	03/07/68 - 03/30/68
		D006072	11-19	04/01/68 - 04/23/68
		D006073	20-26	06/13/68 - 06/29/68
		D006074	27-38	07/01/68 - 07/30/68 (a)
		D006075	39-50	08/02/68 - 08/31/68 (b)
		D006076	51-61	09/02/68 - 09/28/68
		D006077	62-73	10/01/68 - 10/30/68
		D006078	74-85	11/01/68 - 11/30/68
		D006079	86-97	12/02/68 - 12/31/68 (c)
		D006080	98-108	01/02/69 - 01/31/69 (d)
DR003946	DS003946	D006081	1-10	02/03/69 - 02/26/69
		D011544	11-22	02/28/69 - 03/29/69 (e)
		D011545	23-33	04/01/69 - 04/30/69
		D011546	34-45	05/02/69 - 05/31/69 (f)

- (a) D006074 - read error occurred in record 31 of file 1.
- (b) D006075 - read error occurred in record 2 of file 6.
- (c) D006079 - read error occurred in record 2 of file 1.
- (d) D006080 - read error occurred in records 22, 26, 30, 32 of file 1.
- (e) D011544 - read error occurred in record 44 of file 12.
- (f) D011546 - read error occurred in record 26 of file 12.

OGD-5

O, HE + H ION CONC + Ephem, Tape
68-014A-18A

DR+DS03945

~~LR10262 NC1023~~, D6071-80, C4591-99, C5345, 6250+3480, BSA
1108, 108 files, 3280 recs, 112283503 Bytes, 36001 Blks, 1
Variable, 154 recs, 3/7/68 - 4/23/68 (F1-19), 4/13/68-1/31/69 (F20-108)
REMARKS: Read error on Rec 31 file 1 of D6074, Rec 2, file 6 of
D6075, Rec 2, file 1 of D6079, Recs 22, 26, 30, 432, File 1 of
D6080, CARLISLE, B26943215

DR+DS03946

~~LR10152 NC1024~~, D6081, D11544-46, C4600, C8905-7, 6250+3480, BSA
1108, 45 files, 1577 recs, 54143561 Bytes, 36001 Blks, 1
Variable, 152 recs, 2/3/69 - 5/31/69, REMARKS: Read error
Rec 44, file 12 on D11544, Rec 26, file 12 on D11546, BASF,
182/2157 July 91 Lot 23

OGO-5 SPECTROMETER

68-014A-18A

This data set consists of 14 tapes. These tapes are 800 BPI, 7-track, binary with anywhere from 9 to 14 files per tape. Note: each tape is ended with a triple EOF.

The tapes contain header records in BCD format. The data records consist of 36 bit floating point words. The format for the data records is:

<u>WORD</u>	<u>DESCRIPTION</u>
1	JULIAN DAY
2	UNIVERSAL TIME
3	ION CONCENTRATION
4	GEODETTIC LATITUDE
5	GEODETTIC LONGITUDE
6	ALTITUDE
7	L
8	DISTANCE
9	LOCAL TIME
10	MAGNETIC LATITUDE
11	EGRESS LATITUDE
12	INGRESS LATITUDE

The time spans for the tapes are:

<u>D#</u>	<u>C#</u>	<u>FILES</u>	<u>TIME SPAN</u>
D-06071	C-04591	12	3/07/68 - 3/30/68
D-06072	C-05345	11	4/01/68 - 4/23/68
D-06073	C-04592	9	6/13/68 - 6/29/68
D-06074	C-04593	14	7/01/68 - 7/30/68
D-06075	C-04594	14	8/02/68 - 8/31/68

<u>D#</u>	<u>C#</u>	<u>FILES</u>	<u>TIME SPAN</u>
D-06076	C-04595	13	9/02/68 - 9/28/68
D-06077	C-04596	14	10/01/68 - 10/30/68
D-06078	C-04597	14	11/01/68 - 11/30/68
D-06079	C-04598	14	12/02/68 - 12/31/68
D-06080	C-04599	13	1/02/69 - 1/31/69
D-06081	C-04600	12	2/03/69 - 2/26/69
D-11544	C-08905	14	2/28/69 - 3/29/69
D-11545	C-08906	13	4/01/69 - 4/30/69
D-11546	C-08907	14	5/02/69 - 5/31/69

Lockheed
PALO ALTO
RESEARCH
LABORATORY

Department 52-12 Building 202

3251 HANOVER STREET • PALO ALTO, CALIFORNIA • 94304

5 May 1971

National Space Science Data Center
Code 601
Goddard Space Flight Center
Greenbelt, Maryland 20771

Gentlemen:

Accompanying this letter is a shipment of processed magnetic data tapes from the experiment E-18, A Light Ion Mass Spectrometer, for OGO-5.

The data contained on these magnetic tapes are in final analyzed form for the data acquired for the first year of instrument operation. This shipment is in accordance with "A Plan for Submission of Data from Experiment E-18 on OGO-V," Contract No. NAS 5-9092, and forms a portion of the data to be submitted to the National Space Science Data Center as per contract specifications. Additional data will be submitted in the form of charts, graphs, tables, etc., which formed the basis of published results. These additional data will accompany or appropriately be included in the final report to be subsequently submitted.

The format of the data contained on the magnetic tapes included herein is as follows: Each magnetic tape is written with 800 bits per inch, odd parity and binary; each word is a 36 bit word floating point, the left nine bits forming the characteristic; the first record is an indicative record written in BCD format, containing the number of files on tape, the month of year of the data acquisition and the beginning and ending orbit numbers, and address of principal investigators; the files are in ordered records of Oxygen (oxy), Helium (Hel) and Hydrogen (Hyd); each record consists of a 12 by X matrix with $X \leq 500$, the matrix elements in order are (1) day 2) universal time 3) ion concentration 4) geodetic latitude 5) geodetic longitude 6) altitude 7) L 8) distance 9) local time 10) magnetic latitude 11) egress latitude 12) ingress latitude. At the end of tape are placed 3 end of file marks.

A complete description of all computer programs developed for data analysis will be included within the final report.

Respectfully,

Kent K. Harris
Dr. Kent K. Harris
OGO-5 Experimentor
Space Sciences Laboratory

Enclosure

Chapter 3

DATA ANALYSIS

3.1 DATA

The data from the LIMS is received from Goddard Space Flight Center (GSFC) in the form of magnetic tapes. The data have been preliminarily processed to be compatible with the Univac 1108 computing facility of Lockheed Missiles and Space Company. The processed tapes have been remarkably free from errors. Only on rare occasions has it been necessary to return a magnetic tape to GSFC to be reprocessed because it contained noisy or otherwise unmanageable data. Along with the basic data tapes magnetic tapes containing a complete set of orbit-altitude information were also supplied by GSFC. The analysis of the data from the LIMS required utilizing both the data tape and the orbit-attitude data as described in the following sub-section.

3.2 THE APPROACH TO ANALYSIS

The primary purpose of the LIMS was to obtain data on the cold ion distribution within the plasmasphere and plasmatrough. Although many others functions could be performed by the LIMS in the OGO-5 mission (e.g. measure the low energy component of the solar wind, measure the thermal H^+ ion distribution in the magnetosheath, determine bow shock, magnetosheath and magnetopause boundaries, etc.), these were relegated to secondary importance and thereby given only casual attention. The majority of the effort in

reducing and analyzing the LIMS data was confined to the region within approximately 12 earth radii (Re).

As has been described earlier in this report the LIMS is an instrument designed to look in the direction of the spacecraft velocity vector. In the region where the spacecraft velocity is large compared to the most probable ion thermal velocity, α , the ions are "scooped in" by the instrument. The measured ion current is then simply related to the ambient cold plasma density by

$$I_1^+ = N_1^+ AV \cos \theta F_1 \quad (3-1)$$

where: I_1^+ = ion current of species 1 measured by the electrometer of the LIMS.

N_1^+ = the ambient density of species 1

A = the area of the entrance aperture of the LIMS

V = velocity of the spacecraft,

θ = angle between the velocity vector and the normal to the entrance aperture plane

F = the product of the transmission of the spectrometer and gain of the electron multiplier for species 1.

The above expression (Eq. 3-1) is accurate to better than 3% even when the most probable ion velocity is equal to the spacecraft velocity. As the most probable ion velocity increases relative to the spacecraft velocity the relation becomes less and less valid. This relation is valid to within a factor of two even when α is as much as a factor of 5 greater than V .

The most probable ion velocity α depends upon the temperature and mass of the ion:

$$\alpha = (2FT/M)^{1/2} \quad (3-2)$$

therefore α is uncertain to the degree that the ion temperature is unknown.

For the region within the plasmasphere (i.e. when the spacecraft is within from 6 to 8 Re) the ion temperature is probably somewhat less than 10,000°C. However, at this temperature and assuming theOGO-5 vehicle velocity corresponding to approximately 5 Re, the error in utilizing Eq. 3-1 for O^+ ions, He^+ ions and H^+ ions is 0%, 9% and 40% respectively. Thus for analysis of the LIMS data, keeping in mind the uncertainty of the ion temperatures, Equation 3-1 will be utilized to obtain ambient densities of the cold plasma.

The ion densities were calculated from Equation 3-1 with angles θ and velocities V determined from the orbit-attitude tape. The concentration data thus calculated were merged with several of the important geophysical parameters from the orbit-attitude tape and a final data magnetic tape was generated. The data for the entire first year of operation were analyzed and final data tapes constructed. The format of the data contained on the magnetic tapes is as follows: Each magnetic tape is written with 800 bits per inch, odd parity and binary; each word is a 36 bit word floating point, the left nine bits forming the characteristic; the first record is an indicative record written in B. C. D. format, containing the number of files on tape, the month of the year of data acquisition, and the beginning and ending orbit numbers, and address of principal investigators; the files are

in ordered records of oxygen (OXY), helium (Hel) and hydrogen (Hyd); each record consists of a 12 by x matrix with $x \leq 500$, the matrix elements in order are 1) day, 2) universal time, 3) ion concentration 4) geodetic latitude, 5) geodetic longitude, 6) altitude, 7) L, 8) distance, 9) local time, 10) magnetic latitude, 11) egress latitude, 12) ingress latitude. At the end of tape are placed 3 end of file marks. Each final data tape contains data for one specific month of instrument operation.

In the region of the magnetosheath the thermal distribution of H^+ ions is readily measured between 0-600 electron volts. The LIMS covers this region for H^+ ions very well but is displayed only when the larger data rates from the telemetry system (8 and 64 kilobits) are being used. Consequently several 0 to 600 eV thermal distributions of H^+ ions have been acquired and analyzed. These analyses are confined roughly within the orbital range from 8 Re outbound to apogee, and from apogee inbound to 8 Re. From these data the crossing of the boundaries of the various magnetospheric regions were readily discernable. The crossing of the bow shock was particularly dramatic. Signal levels from the LIMS dropped as much as 3 orders of magnitude when passing from the magnetosheath region through the bow shock and out into the solar wind.

The proton energy spectrum is determined for the energy range 0-600 volts in the period of time of the basic sweep cycle of the LIMS, i.e. approx. 4 sec.

The LIMS was designed to focus cold (essentially 0 volt) protons (H^+) through the instrument when the accelerating voltage was -600 volts. A proton with 10 volts energy and properly directed into the entrance aperture of the instrument will focus at -590 volts. In like manner, a 600 volt proton directed with 0° angle of attack into the instrument will focus through

the instrument when the accelerating voltage is 0 volts. In this latter case, the proton signal will appear at the beginning of what is called the "oxygen" state of the instrument. It is then clear that the energy of an entering ion is given by

$$E_p = 600 + V_s \quad (3-3)$$

where E_p = proton energy in volts
 V_s = sweep potential in volts.

Rather than detect a peak as in the case of the normal mass spectroscopy mode, the signal is then spread out over the full range of the sweep due to the energy distribution of the protons that are being focused thru the instrument. The detailed shape of the sweep of the accelerating potential as a function of time is known. With this information the signal out of the electrometer is converted from a time dependent function to an accelerating potential dependent function. This, in turn, is then converted to the proton energy distribution. The display of this data will then be in the form of the proton energy distribution for specific universal times. These times being the start of each sweep cycle.

3.3 DATA ANALYSIS PROBLEMS

Many circumstances have occurred which have complicated the analysis of the data. Telemetry points on equipment group #2 were shared with experiment E-26 under a pre-arranged schedule. The data from the spacecraft when being transmitted on equipment group #2 therefore frequently came in piecemeal form. This substantially complicated the data analysis procedures; but even more important, it frequently created discontinuities in the LIMS

data, sometimes at rather critical points. The problem of piecemeal acquisition of data was compounded further by the failure of wideband transmitter A. The data transmittal procedures utilized after this failure returned the data for relatively short periods of operation.

The electron multiplier in the LIMS, as has been previously described, operated serially and cyclic in two modes, a high gain mode and a low gain mode. It was discovered that calibration of the high gain mode had changed relative to the low gain mode. It was suspected that this situation did arise from a small deposit of vacuum pump oil on the multiplier plates during the extensive pre-launch testing of the instrument. The high gain calibration was then corrected by normalizing to the low gain calibration in their region of overlap. This circumstance continued from initial turn on until about March 20, 1968, when the gain of the high gain mode shifted back to nearly the value of laboratory calibration. This phenomenon was interpreted as an expurgation of the surface of the multiplier plates by the removal of the vacuum pump oil. These two operational characteristics of the LIMS were appropriately compensated for in the computer programs that were developed to analyze the data. After these initial problems with the electron multiplier, no further abrupt changes in multiplier operation were detected except perhaps for the ultimate failure of the multiplier power supply. This can not be properly determined until the second year of data is analyzed. The inflight calibration system of the LIMS showed that a slight multiplier decay transpired throughout much of the flight. This decay was carefully followed and the final data were corrected to reflect the slightly changing gain of the multiplier.

One further difficulty in data analysis occurred when the sweep monitor circuit failed. As described above, the entire mode of operation

of the LIMS assumed a new pattern. This of course required another programming change to properly extract the data.

3.4 COMPUTER PROGRAMS

The computer programs that have been developed to analyze the LIMS data fall basically into two categories. One category encompasses the primary program of data analysis, i.e., the cold plasma in the region of the plasmasphere and plasma trough. The other category is that of the proton energy distribution in the 0-600 eV range. In the first category 3 similar computer programs were developed to reduce the data to an analyzable form. The three programs correspond to the 3 modes of instrument operation described above. The first program, OGO-E-1, is utilized for the data prior to 20 March 1968. This program compensates for the shift in calibration of the high gain mode of the electron multiplier. The second program, OGO-E-2, is utilized for data after 20 March 1968 to 22 April 1968. During this period the multiplier calibration was re-established near the laboratory level. After 22 April 1968 the third computer program OGO-E-3 was used. This program incorporated the change in the instrument logic caused by the failure of the sweep monitor circuit. In addition to these 3 basic programs, other computer programs were developed to display the data in convenient forms. The most frequently used program was that program which plotted the ion concentrations as a function of the geomagnetic parameter, L. The listings of this plot program, the ion concentration programs OGO-E-1 and OGO-E-3 and the proton energy distribution from 0-600 eV are given in the Appendices.

These magnetic tapes have been received by the National Space
Science Data Center.

OGO-5, Experiment E-18

Tape #

OGO-5	March 68
OGO-5	April 68
OGO-5	June 68
OGO-5	July 68
OGO-5	August 68
OGO-5	September 68
OGO-5	October 68
OGO-5	November 68
OGO-5	December 68
OGO-5	January 69
OGO-5	February 69

Signed

Return to: Dr. Kent K. Harris
Dept. 52-12, Bldg. 202
Lockheed Palo Alto Research Lab.
3251 Hanover Street
Palo Alto, California 94304

2225	+167321075255	+206419334583	02227	+206417527200	+220506132207
2229	+211776764734	+206543606557	02231	-170367424555	+207430124630
2233	+206710000000	+220663673561	02235	+172721206174	+206537377073
2237	+167321077441	+220415421044	02239	+204617623325	+220500220471
2241	+211777640530	+206543601623	02243	-170367417770	+207430130044
2245	+206710000000	+220663716116	02247	+173416173624	+206537370120
2249	+167321101615	+220416507324	02251	+204617717452	+220500306752
2253	+211777114323	+206543574667	02255	-170367413201	+207430133257
2257	+206710000000	+220663740453	02259	+172721444105	+206537361142
2261	+167321103765	+220416575602	02263	+204620013564	+220500375232
2265	+211777170110	+206543567726	02267	-170367406413	+207430130473
2269	+206710000000	+220663763011	02271	+172721562736	+206537352052
2273	+167321102777	+220416633777	02275	+204620107134	+220500463430
2277	+211777243401	+206543562542	02279	-170367401743	+207430141727
2281	+206710000000	+220664000655	02283	+000000000000	+000000000000
		WORDS 02285 TO 02292 ARE ALL +000000000000			
2293	+206710000000	+220664032240	02295	+172722137240	+206537324401
2297	+167321114033	+220417126855	02299	+204620371413	+222500726211
2301	+211777445446	+206543543006	02303	-170367364145	+207430133653
2305	+206710000000	+220664074578	02307	+173416616305	+206537315511
2309	+167321110043	+220417214747	02311	+204620464760	+220501014404
2313	+211777320736	+206543535622	02315	-170367357473	+207430157107
2317	+206710000000	+220664117132	02319	+172722374714	+206537306420
2321	+167321120055	+220417303141	02323	+204620560325	+220501102600
2325	+211777574225	+206543530437	02327	-170367353021	+207430162343
2329	+206710000000	+220664141473	02331	+172607303477	+206537277330
2333	+167321122067	+220417371335	02335	+204620653675	+220501170775
2337	+211777647516	+206543533253	02339	-170367346347	+207430165577
2341	+206710000000	+220664164025	02343	+173417027510	+206537270237
2345	+167321124101	+220417457527	02347	+204620747242	+220501267171
2349	+211777722005	+206543516067	02351	-170367341675	+207430171032
2353	+206710000000	+220664206362	02355	+172722751222	+206537261147
2357	+167321120111	+220417545721	02359	+204621042607	+220501345364
2361	+211777720378	+206543510700	02363	-170367336223	+207430174256

IBTD of last record
of last tape in data
set to check steps from
against 360/30 dump.

2365	+20671000000	+206664230717	02367	+172507611677	+206537252657
2369	-167321130123	+206417634113	02371	+206621136154	+220501433557
2373	+212400024671	+206543503517	02375	-170367330551	+207430177528
2377	+20671000000	+220664253254	02379	+172723206676	+206537242767
2381	-167321132135	+2206417722305	02383	+204621231521	+220501521752
2385	+212400056425	+206543476133	02387	-170367324077	+207430262756
2389	+20671000000	+220664275612	02391	+173417316514	+206537233576
2393	-167321134145	+220420010502	02395	+204621325071	+220501610150
2397	+212400100162	+206543471147	02399	-170367317425	+207430262112
2401	+20671000000	+220664320147	02403	+173417374314	+206537224605
2405	-167321136157	+220420076673	02407	+204621420432	+220501676342
2409	+212400125713	+206543663762	02411	-170367312754	+207430211446
2413	+20671000000	+220664343554	02415	+172610222055	+206537215402
2417	-167321140011	+220420164775	02419	+204621513213	+220501764447
2421	+212400152271	+206543456336	02423	-170367306427	+207430214727
2425	+20671000000	+220664365041	02427	+172410324025	+206537206177
2429	-167321141545	+220420253102	02431	+204621605774	+220502053553
2433	+212400200646	+206543450712	02435	-170367302102	+207430220207
2437	+20671000000	+220664407376	02439	+172610425777	+206537176774
2441	-167321143477	+220420341205	02443	+204621700553	+220502140560
2445	+212400226284	+206543443267	02447	-170367275555	+207430223467
2449	+20671000000	+220664431733	02451	+173417661026	+206537167570
2453	-167321145333	+220420427310	02455	+204621773330	+220502226765
2457	+212400253601	+206543435643	02459	-170367271227	+207430226750
2461	+20671000000	+220664454271	02463	+172724255524	+206537150365
2465	-167321147167	+220420515017	02467	+204622066122	+220502315075
2469	+212400301160	+206543430217	02471	-170367264792	+207430232230
2473	+20671000000	+220664476626	02475	+172724374232	+206537151162
2477	-167321151021	+220420603522	02479	+204622160703	+220502403201
2481	+212400326525	+206543422873	02483	-170367263355	+207430235511

D-6071

0001	2414244156057	211614153105	152423052205	333035302512	131327242212	311227057514	340530150627	250532514125
0049	202005105067	271613005505	241020151212	110525062124	050621312405	271230120627	101505210670	050505050505
0097	151516300301	052512051024	233105162330	051106310605	132427053115	120522242331	150521430522	252171150511
0145	156620050524	270716313005	510541156151			510541156151	510541156151	510541156151

0001	207614000000	214423000202	176533713210	233577565325	210470722357	220614415541	204435060170	220676255353
0049	213405711217	576374272644	571047631254	227454062343	207414000000	214423103702	000000000000	000000000000
0097	000000000000	000000000000	000000000000	000000000000	000000000000	000000000000	000000000000	000000000000
0145	207414000000	214425233715	175536774651	233576764351	210470630065	220614105135	204435260533	220675754423
0193	213406723327	576376235577	571047713252	227454062242	207414000000	214426351563	175711542313	220675463257
0241	210470572760	220513740500	204435160651	220756510055	213406730407	576344615574	571047744314	207454050214
0289	207414000000	214427467431	175524427255	233676161646	210470535657	220613574150	204435060757	220675413655
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5665	567375646675	217557347451	167657576347	206527770323	567374613037	217557530244	204511277117	206530013171
5713	214056000000	206576590031	570459761455	217475215553	214056000000	217475564723	167650494142	206530571715
5761	214407003655	206576760556	204507576115	217723312735	214407016232	206576660340	204512060617	217723473462
5809	567374661203	217557711043	170611753054	2065300075163	567374667143	217556716104	17157174641	206530136607
5857	214056000000	217476631555	570360446055	207402542741	214056000000	217476676540	20451252525	207402583831
5905	214407055344	206577103442	570357475682	206530254037	567374673061	204512601754	171720437076	206530242975
5953	567374770535	217550613605	204512742470	217724215417	214407043046	20657717010274	167650723242	206530525327
5997	214056000000	206577103442	171572214231	206530303605	567375016411	2175607747352	217474320321	217724376146
6045	214407055344	206577103442	570457426714	207402605657	214407114442	206577326203	570357103336	207402617222
6093	567375044261	217551165076	204513473465	217724556657	214407114442	206577326203	570357103336	207402617222

6337	210456000000	217477166757	171440162605	206030607013	567375072131	217561335623	204513704461	217724737370
6395	214407126740	206577407117	570357357761	20402630165	210456000000	217477231662	167660262024	206530450517
6433	567375120003	217561516353	204518145292	217725120103	214407141237	206577470033	570357336402	207402641124
6481	210456000000	217477010544	167590267740	206530512222	567375145653	217561671100	204514405716	217725100614
6525	214407153536	206577550745	570357111925	207402652063	210456000000	217477345437	167660275655	206530553726
6577	567375173523	217562057630	204514646436	217725461127	214407156035	206577361661	570357265447	207402663023
6625	210456000000	217477412331	170514263331	206530615431	567375221375	217562240355	204519107152	217725642040
6672	214407200333	206577712573	570357242371	207402673762	210456000000	217477457223	171440206331	206530657134
6721	567375247245	217562421102	204315347666	217726022890	21440712632	206577773505	570357216514	207402704721
6765	210456000000	217477524115	167560317421	206530720637	567375275115	217562601632	204515610406	217726203264
6817	214407225131	206600054421	570457173135	207402715660	210456000000	217477566554	167660304451	206530757500
6865	567375321311	217562753277	204516036722	217726354716	214407236657	206600313176	570357151130	207402726707
6913	210456000000	217477633447	167560257557	206531020061	567375346671	217563133756	204516277162	217726535361
6961	214407251109	206600210764	570357126060	207402740722	210456000000	217477702575	170614214603	206531052461
7005	567375375531	217563123457	204516551433	217726725044	214407264031	206600273327	570357101645	207402753435
7057	210456000000	217477747467	171440130354	206531123042	567375423111	217563504133	204517011671	217727105504
7105	214407276252	206600352713	570357056575	207402765450	210456000000	217500014361	170614146405	206531163423
7153	567375450467	217563664607	204517252127	217727266144	214407310472	206600432277	570357033525	207402777462
7201	210456000000	217500061254	171440074512	206531224005	567375476047	217564045266	204517512371	217727446607
7245	214407322713	206600511665	570357010454	207403011475	210456000000	217500126146	171440056571	206531264365
7297	567375523427	217564225742	204517782626	217727627247	214407335133	206600571251	570356765404	207403023507
7345	210456000000	217500173041	170514055110	206531324747	567375551007	217564406422	204520213070	217730007712
7393	214407347354	206600650636	570356742333	207403035522	210456000000	217500237733	171440022730	206531365330
7441	567375576367	217564567076	204520453326	217730173352	214407361575	206600730222	570356717263	207403047534
7485	210456000000	217507304625						

Data Set Catalog #143

OGO-5, Spectrometer

68-014A-18A 14 tapes

OGO 5

O, HE + H ION CONCS + EPHEM, TAPE

68-014A-18A

THIS DATA SET HAS BEEN RESTORED. THERE WERE ORIGINALLY FOURTEEN 7-TRACK, 800 BPI TAPES, WRITTEN IN BINARY. THERE ARE TWO RESTORED TAPES, WHICH WERE PADDED DURING THE RESTORATION PROCESS. THE DR TAPES ARE 3480 CARTRIDGES AND THE DS TAPES ARE 9-TRACK, 6250 BPI. THE ORIGINAL TAPES WERE CREATED ON A 1108 COMPUTER. THE DR AND DS NUMBERS ALONG WITH THE CORRESPONDING D NUMBERS AND TIME SPANS ARE AS FOLLOWS:

DR#	DS#	D#	FILES	TIME SPAN
DR003945	DS003945	D006071	1-10	03/07/68 - 03/30/68
		D006072	11-19	04/01/68 - 04/23/68
		D006073	20-26	06/13/68 - 06/29/68
		D006074	27-38	07/01/68 - 07/30/68 (a)
		D006075	39-50	08/02/68 - 08/31/68 (b)
		D006076	51-61	09/02/68 - 09/28/68
		D006077	62-73	10/01/68 - 10/30/68
		D006078	74-85	11/01/68 - 11/30/68
		D006079	86-97	12/02/68 - 12/31/68 (c)
		D006080	98-108	01/02/69 - 01/31/69 (d)
DR003946	DS003946	D006081	1-10	02/03/69 - 02/26/69
		D011544	11-22	02/28/69 - 03/29/69 (e)
		D011545	23-33	04/01/69 - 04/30/69
		D011546	34-45	05/02/69 - 05/31/69 (f)

- (a) D006074 - read error occurred in record 31 of file 1.
 (b) D006075 - read error occurred in record 2 of file 6.
 (c) D006079 - read error occurred in record 2 of file 1.
 (d) D006080 - read error occurred in records 22, 26, 30, 32 of file 1.
 (e) D011544 - read error occurred in record 44 of file 12.
 (f) D011546 - read error occurred in record 26 of file 12.

OGD-5

O, HE + H ION CONC + Ephem, Tape
68-014A-18A

DR+DS03945

~~LR10262 NC1023~~, D6071-80, C4591-99, C5345, 6250+3480, BSA
1108, 108 files, 3280 recs, 112283503 Bytes, 36001 Blks, 1
Variable, 154 recs, 3/7/68 - 4/23/68 (F1-19), 4/13/68-1/31/69 (F20-108)
REMARKS: Read error on Rec 31 file 1 of D6074, Rec 2, file 6 of
D6075, Rec 2, file 1 of D6079, Recs 22, 26, 30, 432, File 1 of
D6080, CARLISLE, B26943215

DR+DS03946

~~LR10152 NC1024~~, D6081, D11544-46, C4600, C8905-7, 6250+3480, BSA
1108, 45 files, 1577 recs, 54143561 Bytes, 36001 Blks, 1
Variable, 152 recs, 2/3/69 - 5/31/69, REMARKS: Read error
Rec 44, file 12 on D11544, Rec 26, file 12 on D11546, BASF,
182/2157 July 91 Lot 23

OGO-5 SPECTROMETER

68-014A-18A

This data set consists of 14 tapes. These tapes are 800 BPI, 7-track, binary with anywhere from 9 to 14 files per tape. Note: each tape is ended with a triple EOF.

The tapes contain header records in BCD format. The data records consist of 36 bit floating point words. The format for the data records is:

<u>WORD</u>	<u>DESCRIPTION</u>
1	JULIAN DAY
2	UNIVERSAL TIME
3	ION CONCENTRATION
4	GEODETTIC LATITUDE
5	GEODETTIC LONGITUDE
6	ALTITUDE
7	L
8	DISTANCE
9	LOCAL TIME
10	MAGNETIC LATITUDE
11	EGRESS LATITUDE
12	INGRESS LATITUDE

The time spans for the tapes are:

<u>D#</u>	<u>C#</u>	<u>FILES</u>	<u>TIME SPAN</u>
D-06071	C-04591	12	3/07/68 - 3/30/68
D-06072	C-05345	11	4/01/68 - 4/23/68
D-06073	C-04592	9	6/13/68 - 6/29/68
D-06074	C-04593	14	7/01/68 - 7/30/68
D-06075	C-04594	14	8/02/68 - 8/31/68

<u>D#</u>	<u>C#</u>	<u>FILES</u>	<u>TIME SPAN</u>
D-06076	C-04595	13	9/02/68 - 9/28/68
D-06077	C-04596	14	10/01/68 - 10/30/68
D-06078	C-04597	14	11/01/68 - 11/30/68
D-06079	C-04598	14	12/02/68 - 12/31/68
D-06080	C-04599	13	1/02/69 - 1/31/69
D-06081	C-04600	12	2/03/69 - 2/26/69
D-11544	C-08905	14	2/28/69 - 3/29/69
D-11545	C-08906	13	4/01/69 - 4/30/69
D-11546	C-08907	14	5/02/69 - 5/31/69

Lockheed
PALO ALTO
RESEARCH
LABORATORY

Department 52-12 Building 202

3251 HANOVER STREET • PALO ALTO, CALIFORNIA • 94304

5 May 1971

National Space Science Data Center
Code 601
Goddard Space Flight Center
Greenbelt, Maryland 20771

Gentlemen:

Accompanying this letter is a shipment of processed magnetic data tapes from the experiment E-18, A Light Ion Mass Spectrometer, for OGO-5.

The data contained on these magnetic tapes are in final analyzed form for the data acquired for the first year of instrument operation. This shipment is in accordance with "A Plan for Submission of Data from Experiment E-18 on OGO-V," Contract No. NAS 5-9092, and forms a portion of the data to be submitted to the National Space Science Data Center as per contract specifications. Additional data will be submitted in the form of charts, graphs, tables, etc., which formed the basis of published results. These additional data will accompany or appropriately be included in the final report to be subsequently submitted.

The format of the data contained on the magnetic tapes included herein is as follows: Each magnetic tape is written with 800 bits per inch, odd parity and binary; each word is a 36 bit word floating point, the left nine bits forming the characteristic; the first record is an indicative record written in BCD format, containing the number of files on tape, the month of year of the data acquisition and the beginning and ending orbit numbers, and address of principal investigators; the files are in ordered records of Oxygen (oxy), Helium (Hel) and Hydrogen (Hyd); each record consists of a 12 by X matrix with $X \leq 500$, the matrix elements in order are (1) day 2) universal time 3) ion concentration 4) geodetic latitude 5) geodetic longitude 6) altitude 7) L 8) distance 9) local time 10) magnetic latitude 11) egress latitude 12) ingress latitude. At the end of tape are placed 3 end of file marks.

A complete description of all computer programs developed for data analysis will be included within the final report.

Respectfully,

Kent K. Harris
Dr. Kent K. Harris
OGO-5 Experimentor
Space Sciences Laboratory

Enclosure

Chapter 3

DATA ANALYSIS

3.1 DATA

The data from the LIMS is received from Goddard Space Flight Center (GSFC) in the form of magnetic tapes. The data have been preliminarily processed to be compatible with the Univac 1108 computing facility of Lockheed Missiles and Space Company. The processed tapes have been remarkably free from errors. Only on rare occasions has it been necessary to return a magnetic tape to GSFC to be reprocessed because it contained noisy or otherwise unmanageable data. Along with the basic data tapes magnetic tapes containing a complete set of orbit-altitude information were also supplied by GSFC. The analysis of the data from the LIMS required utilizing both the data tape and the orbit-attitude data as described in the following sub-section.

3.2 THE APPROACH TO ANALYSIS

The primary purpose of the LIMS was to obtain data on the cold ion distribution within the plasmasphere and plasmatrough. Although many others functions could be performed by the LIMS in the OGO-5 mission (e.g. measure the low energy component of the solar wind, measure the thermal H^+ ion distribution in the magnetosheath, determine bow shock, magnetosheath and magnetopause boundaries, etc.), these were relegated to secondary importance and thereby given only casual attention. The majority of the effort in

reducing and analyzing the LIMS data was confined to the region within approximately 12 earth radii (R_E).

As has been described earlier in this report the LIMS is an instrument designed to look in the direction of the spacecraft velocity vector. In the region where the spacecraft velocity is large compared to the most probable ion thermal velocity, α , the ions are "scooped in" by the instrument. The measured ion current is then simply related to the ambient cold plasma density by

$$I_1^+ = N_1^+ A V \cos \theta F_1 \quad (3-1)$$

where: I_1^+ = ion current of species 1 measured by the electrometer of the LIMS.

N_1^+ = the ambient density of species 1

A = the area of the entrance aperture of the LIMS

V = velocity of the spacecraft,

θ = angle between the velocity vector and the normal to the entrance aperture plane

F = the product of the transmission of the spectrometer and gain of the electron multiplier for species 1.

The above expression (Eq. 3-1) is accurate to better than 3% even when the most probable ion velocity is equal to the spacecraft velocity. As the most probable ion velocity increases relative to the spacecraft velocity the relation becomes less and less valid. This relation is valid to within a factor of two even when α is as much as a factor of 5 greater than V .

The most probable ion velocity α depends upon the temperature and mass of the ion:

$$\alpha = (2FT/M)^{1/2} \quad (3-2)$$

therefore α is uncertain to the degree that the ion temperature is unknown.

For the region within the plasmasphere (i.e. when the spacecraft is within from 6 to 8 Re) the ion temperature is probably somewhat less than 10,000°C. However, at this temperature and assuming theOGO-5 vehicle velocity corresponding to approximately 5 Re, the error in utilizing Eq. 3-1 for O^+ ions, He^+ ions and H^+ ions is 0%, 9% and 40% respectively. Thus for analysis of the LIMS data, keeping in mind the uncertainty of the ion temperatures, Equation 3-1 will be utilized to obtain ambient densities of the cold plasma.

The ion densities were calculated from Equation 3-1 with angles θ and velocities V determined from the orbit-attitude tape. The concentration data thus calculated were merged with several of the important geophysical parameters from the orbit-attitude tape and a final data magnetic tape was generated. The data for the entire first year of operation were analyzed and final data tapes constructed. The format of the data contained on the magnetic tapes is as follows: Each magnetic tape is written with 800 bits per inch, odd parity and binary; each word is a 36 bit word floating point, the left nine bits forming the characteristic; the first record is an indicative record written in B. C. D. format, containing the number of files on tape, the month of the year of data acquisition, and the beginning and ending orbit numbers, and address of principal investigators; the files are

in ordered records of oxygen (OXY), helium (Hel) and hydrogen (Hyd); each record consists of a 12 by x matrix with $x \leq 500$, the matrix elements in order are 1) day, 2) universal time, 3) ion concentration 4) geodetic latitude, 5) geodetic longitude, 6) altitude, 7) L, 8) distance, 9) local time, 10) magnetic latitude, 11) egress latitude, 12) ingress latitude. At the end of tape are placed 3 end of file marks. Each final data tape contains data for one specific month of instrument operation.

In the region of the magnetosheath the thermal distribution of H^+ ions is readily measured between 0-600 electron volts. The LIMS covers this region for H^+ ions very well but is displayed only when the larger data rates from the telemetry system (8 and 64 kilobits) are being used. Consequently several 0 to 600 eV thermal distributions of H^+ ions have been acquired and analyzed. These analyses are confined roughly within the orbital range from 8 Re outbound to apogee, and from apogee inbound to 8 Re. From these data the crossing of the boundaries of the various magnetospheric regions were readily discernable. The crossing of the bow shock was particularly dramatic. Signal levels from the LIMS dropped as much as 3 orders of magnitude when passing from the magnetosheath region through the bow shock and out into the solar wind.

The proton energy spectrum is determined for the energy range 0-600 volts in the period of time of the basic sweep cycle of the LIMS, i.e. approx. 4 sec.

The LIMS was designed to focus cold (essentially 0 volt) protons (H^+) through the instrument when the accelerating voltage was -600 volts. A proton with 10 volts energy and properly directed into the entrance aperture of the instrument will focus at -590 volts. In like manner, a 600 volt proton directed with 0° angle of attack into the instrument will focus through

the instrument when the accelerating voltage is 0 volts. In this latter case, the proton signal will appear at the beginning of what is called the "oxygen" state of the instrument. It is then clear that the energy of an entering ion is given by

$$E_p = 600 + V_s \quad (3-3)$$

where E_p = proton energy in volts
 V_s = sweep potential in volts.

Rather than detect a peak as in the case of the normal mass spectroscopy mode, the signal is then spread out over the full range of the sweep due to the energy distribution of the protons that are being focused thru the instrument. The detailed shape of the sweep of the accelerating potential as a function of time is known. With this information the signal out of the electrometer is converted from a time dependent function to an accelerating potential dependent function. This, in turn, is then converted to the proton energy distribution. The display of this data will then be in the form of the proton energy distribution for specific universal times. These times being the start of each sweep cycle.

3.3 DATA ANALYSIS PROBLEMS

Many circumstances have occurred which have complicated the analysis of the data. Telemetry points on equipment group #2 were shared with experiment E-26 under a pre-arranged schedule. The data from the spacecraft when being transmitted on equipment group #2 therefore frequently came in piecemeal form. This substantially complicated the data analysis procedures; but even more important, it frequently created discontinuities in the LIMS

data, sometimes at rather critical points. The problem of piecemeal acquisition of data was compounded further by the failure of wideband transmitter A. The data transmittal procedures utilized after this failure returned the data for relatively short periods of operation.

The electron multiplier in the LIMS, as has been previously described, operated serially and cyclic in two modes, a high gain mode and a low gain mode. It was discovered that calibration of the high gain mode had changed relative to the low gain mode. It was suspected that this situation did arise from a small deposit of vacuum pump oil on the multiplier plates during the extensive pre-launch testing of the instrument. The high gain calibration was then corrected by normalizing to the low gain calibration in their region of overlap. This circumstance continued from initial turn on until about March 20, 1968, when the gain of the high gain mode shifted back to nearly the value of laboratory calibration. This phenomenon was interpreted as an expurgation of the surface of the multiplier plates by the removal of the vacuum pump oil. These two operational characteristics of the LIMS were appropriately compensated for in the computer programs that were developed to analyze the data. After these initial problems with the electron multiplier, no further abrupt changes in multiplier operation were detected except perhaps for the ultimate failure of the multiplier power supply. This can not be properly determined until the second year of data is analyzed. The inflight calibration system of the LIMS showed that a slight multiplier decay transpired throughout much of the flight. This decay was carefully followed and the final data were corrected to reflect the slightly changing gain of the multiplier.

One further difficulty in data analysis occurred when the sweep monitor circuit failed. As described above, the entire mode of operation

of the LIMS assumed a new pattern. This of course required another programming change to properly extract the data.

3.4 COMPUTER PROGRAMS

The computer programs that have been developed to analyze the LIMS data fall basically into two categories. One category encompasses the primary program of data analysis, i.e., the cold plasma in the region of the plasmasphere and plasma trough. The other category is that of the proton energy distribution in the 0-600 eV range. In the first category 3 similar computer programs were developed to reduce the data to an analyzable form. The three programs correspond to the 3 modes of instrument operation described above. The first program, OGO-E-1, is utilized for the data prior to 20 March 1968. This program compensates for the shift in calibration of the high gain mode of the electron multiplier. The second program, OGO-E-2, is utilized for data after 20 March 1968 to 22 April 1968. During this period the multiplier calibration was re-established near the laboratory level. After 22 April 1968 the third computer program OGO-E-3 was used. This program incorporated the change in the instrument logic caused by the failure of the sweep monitor circuit. In addition to these 3 basic programs, other computer programs were developed to display the data in convenient forms. The most frequently used program was that program which plotted the ion concentrations as a function of the geomagnetic parameter, L. The listings of this plot program, the ion concentration programs OGO-E-1 and OGO-E-3 and the proton energy distribution from 0-600 eV are given in the Appendices.

These magnetic tapes have been received by the National Space
Science Data Center.

OGO-5, Experiment E-18

Tape #

OGO-5	March 68
OGO-5	April 68
OGO-5	June 68
OGO-5	July 68
OGO-5	August 68
OGO-5	September 68
OGO-5	October 68
OGO-5	November 68
OGO-5	December 68
OGO-5	January 69
OGO-5	February 69

Signed

Return to: Dr. Kent K. Harris
Dept. 52-12, Bldg. 202
Lockheed Palo Alto Research Lab.
3251 Hanover Street
Palo Alto, California 94304

2225	+167321075255	+206419334583	02227	+206417527200	+220506132207
2229	+211776764734	+206543606557	02231	-170367424555	+207430124630
2233	+206710000000	+220663673561	02235	+172721206174	+206537377073
2237	+167321077441	+220415421044	02239	+204617623325	+220500220471
2241	+211777640530	+206543601623	02243	-170367417770	+207430130044
2245	+206710000000	+220663716116	02247	+173416173624	+206537370120
2249	+167321101615	+220416507324	02251	+204617717452	+220500306752
2253	+211777114323	+206543574667	02255	-170367413201	+207430133257
2257	+206710000000	+220663740453	02259	+172721444105	+206537361142
2261	+167321103765	+220416575602	02263	+204620013564	+220500375232
2265	+211777170110	+206543567726	02267	-170367406413	+207430130473
2269	+206710000000	+220663763011	02271	+172721562736	+206537352052
2273	+167321102777	+220416633777	02275	+204620107134	+220500463430
2277	+211777243401	+206543562542	02279	-170367401743	+207430141727
2281	+206710000000	+220664000655	02283	+200000000000	+200000000000
		WORDS 02285 TO 02292 ARE ALL +000000000000			
2293	+206710000000	+220664032240	02295	+172722137240	+206537324401
2297	+167321114033	+220417126855	02299	+204620371413	+222500726211
2301	+211777445446	+206543543006	02303	-170367364145	+207430133653
2305	+206710000000	+220664074578	02307	+173416616305	+206537315511
2309	+167321116043	+220417214747	02311	+204620464760	+220501014404
2313	+211777320736	+206543535622	02315	-170367357473	+207430157107
2317	+206710000000	+220664117132	02319	+172722374714	+206537306420
2321	+167321120055	+220417303141	02323	+204620560325	+220501102600
2325	+211777574225	+206543530437	02327	-170367353021	+207430162343
2329	+206710000000	+220664141473	02331	+172607303477	+206537277330
2333	+167321122067	+220417371335	02335	+204620653675	+220501170775
2337	+211777647516	+206543533253	02339	-170367346347	+207430165577
2341	+206710000000	+220664164025	02343	+173417027510	+206537270237
2345	+167321124101	+220417457527	02347	+204620747242	+220501267171
2349	+211777722005	+206543516067	02351	-170367341675	+207430171032
2353	+206710000000	+220664206362	02355	+172722751222	+206537261147
2357	+167321126111	+220417545721	02359	+204621042607	+220501345364
2361	+211777720378	+206543510700	02363	-170367336223	+207430174256

IBTD of last record
of last tape in data
set to check steps from
against 360/30 dump.

2365	+20671000000	+206664230717	02367	+172507611677	+206537252657
2369	-167321130123	+206417634113	02371	+206621136154	+220501433557
2373	+212400024671	+206543503517	02375	-170367330551	+207430177528
2377	+20671000000	+220664253254	02379	+172723206676	+206537242767
2381	-167321132135	+2206417722305	02383	+204621231521	+220501521752
2385	+212400056425	+206543476133	02387	-170367324077	+207430262756
2389	+20671000000	+220664275612	02391	+173417316514	+206537233576
2393	-167321134145	+220420010502	02395	+204621325071	+220501610150
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D-6071

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