

DATA SET CATALOG #47

Extraterrestrial solar spectral irradiance
at
earth's mean solar distance

D-4899

CORR551

1 tape

Table of Contents

1. Introduction
2. Errata/Change Log
3. LINKS TO RELEVANT INFORMATION IN THE ONLINE NSSDC
INFORMATION SYSTEM
4. Catalog Materials
 - a. Associated Documents
 - b. Core Catalog Materials

1. INTRODUCTION:

The documentation for this data set was originally on paper, kept in NSSDC's Data Set Catalogs (DSCs). The paper documentation in the Data Set Catalogs have been made into digital images, and then collected into a single PDF file for each Data Set Catalog. The inventory information in these DSCs is current as of July 1, 2004. This inventory information is now no longer maintained in the DSCs, but is now managed in the inventory part of the NSSDC information system. The information existing in the DSCs is now not needed for locating the data files, but we did not remove that inventory information.

The offline tape datasets have now been migrated from the original magnetic tape to Archival Information Packages (AIP's).

A prior restoration may have been done on data sets, if a requestor of this data set has questions; they should send an inquiry to the request office to see if additional information exists.

2. ERRATA/CHANGE LOG:

NOTE: Changes are made in a text box, and will show up that way when displayed on screen with a PDF reader.

When printing, special settings may be required to make the text box appear on the printed output.

Version	Date	Person	Page	Description of Change
01				
02				

3 LINKS TO RELEVANT INFORMATION IN THE ONLINE NSSDC
INFORMATION SYSTEM:

<http://nssdc.gsfc.nasa.gov/nmc/>

[NOTE: This link will take you to the main page of the NSSDC Master Catalog. There you will be able to perform searches to find additional information]

4. CATALOG MATERIALS:

- a. Associated Documents To find associated documents you will need to know the document ID number and then click here.
<http://nssdcftp.gsfc.nasa.gov/miscellaneous/documents/>

- b. Core Catalog Materials

REQ. AGENT

DAD

ACQ. AGENT

JRJ

EXTRATERRESTRIAL SOLAR SPECTRAL IRRADIANCE

MS-12A

This data set is contained on 1 file on a 9-track, 1600 BPI, ASCII magnetic tape created on the MODCOMP IV computer from card images. The file consists of solar spectral irradiance data covering the wavelength range 300 - 2500 nanometers. This data was originally created at 556 BPI, 7-track BCD on the IBM 7094 computer but was converted to 9-track ASCII because of the phasing out of 7-track tape drives. The document number is B-04225-000A. The D and C numbers follow below:

D#

D-04899

C#

C-02555

EXTRATERRESTRIAL SOLAR SPECTRAL IRRADIANCE TAPE FORMAT

RECORDS 1 - 15
COLUMNS

1 - 76 Alphameric Data for Title Page

77 - 80 Sequence Number (Integer)

RECORDS 16 - 19
COLUMNS

1 - 76 Alphameric Data for Page Headings

77 - 80 Sequence Number (Integer)

RECORDS 20 - N
COLUMNS

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12 - 18 Wave Length, Nanometer (F7.1)

19 - 31 Blank

32 - 38 Extraterrestrial Spectral Irradiance,
Watt/Sqmeter - Nanometer (F7.3)

39 - 55 Blank

56 - 62 Percentage of Total Irradiance at
Shorter Wave Length (F7.2)

63 - 76 Blank

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B04225-000A

PBIN = 80

GB002A-01

DSC 47

Determination of Extraterrestrial Solar Spectral Irradiance from a Research Aircraft

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and B. Douglas Pearson, Jr.***

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Determination of Extraterrestrial Solar Spectral Irradiance from a Research Aircraft

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Results are presented of an experiment to determine extraterrestrial solar spectral irradiance at the earth's mean solar distance within the 300–2500 nm wavelength region. Spectroradiometric measurements were performed during eleven research flights on board a NASA CV-990 aircraft at altitudes between 11.6 km and 12.5 km. Precision of the measurements was better than $\pm 1\%$. Absolute accuracy of the resultant extraterrestrial solar spectral irradiance is approximately $\pm 3\%$ over most of the measurement range. A listing of results is presented at intervals varying from 0.1 nm throughout most of the uv-visible Fraunhofer region to 5 nm in the continuum region of the ir. Additionally, a listing of solar spectral irradiance, smoothed over the detailed Fraunhofer structure, is presented for engineering use.

I. Introduction

The National Aeronautics and Space Administration (NASA) has a strong interest in a more accurate determination of the total and spectral intensity of sunlight in space. This interest stems from the fact that absorbed solar radiation constitutes the primary source of power and heat for spacecraft. In addition, the degrading influence of solar uv radiation on materials used in spacecraft construction requires that this important spectral region be known to a higher accuracy than at present. Research currently under way at NASA in solar physics, terrestrial and atmospheric physics, and meteorology also requires accurate knowledge of solar spectral irradiance at the top of the earth's atmosphere.

The present knowledge of extraterrestrial solar spectral irradiance has been obtained primarily through ground-based measurements from various locations around the world during a period of over fifty years. In performing these measurements, the observer has been severely handicapped by atmospheric opacity and variability due to the presence of smoke, haze, dust, water vapor, and other constituents found near ground

level. The extrapolation to outside the atmosphere involves performing measurements through different amounts of atmosphere corresponding to different times of day; thus, variable transmittance significantly affects the accuracy of the results.

A way to partially circumvent the problems encountered in ground-based experiments is to perform measurements from an altitude where atmospheric transparency and constancy allow accurate extrapolation to extraterrestrial conditions. A platform suitable for these measurements is the CV-990 jet aircraft (commercial designation: Convair CV-990) owned by NASA and specially modified for use in astronomical and space sciences research.¹

Recent advances in spectroradiometry, including the establishment of a Standard of Spectral Irradiance by the National Bureau of Standards² and an improved understanding of subtle errors in spectral irradiance measurements,³ have made possible significant improvements in instrumentation with resultant improvements in accuracy. The improved instrumentation, coupled with the ability to perform measurements from altitudes where variable components, such as water vapor and aerosol, are reduced to an insignificant level, gave promise that a significant reduction in the uncertainty of solar spectral irradiance could be achieved. To this end, an experiment was designed and flown on a series of eleven flights in the summer and autumn of 1967 to perform solar spectral irradiance measurements as part of a total effort directed toward reducing present uncertainties in total and spectral solar irradiance. This paper presents the most significant results of that experiment.

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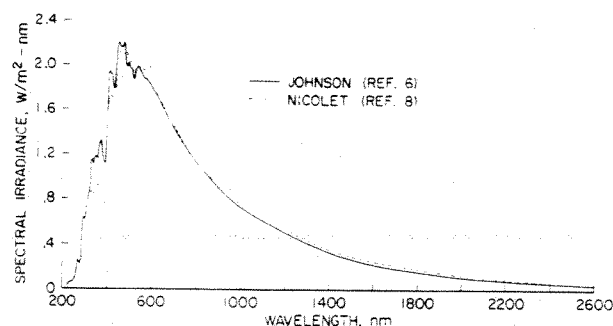


Fig. 1. Comparison of solar spectral irradiance, Johnson and Nicolet.

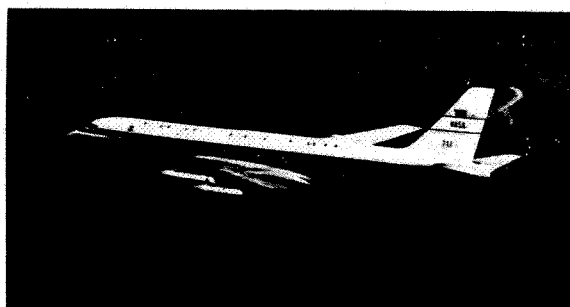


Fig. 2. NASA CV-990 jet aircraft showing upper and side observation windows.

II. Present Knowledge of Solar Spectral Irradiance

The absolute spectral intensity of radiation from the sun has been obtained primarily through ground-based measurements. Rocket spectrograph measurements have been made in uv spectral regions inaccessible to earth-based observations. However, because of the lack of adequate spectral standards in the uv below 250 nm, these measurements were performed on a relative basis and adjusted to ground-based data in a region of overlap. When a comparison of existing measurements is made, it is apparent that discrepancies, especially in the uv below 400 nm, exist as large as 20–30%. In addition, most of the measurements on an absolute energy scale are integrated over 10-nm bands. While this resolution is adequate for certain purposes, a measurement of spectral irradiance on an absolute scale with greatly improved resolution would be beneficial in many areas.

Several excellent reviews of the present state of knowledge of solar spectral irradiance in the 200–7000 nm wavelength range are available in the literature.^{4,5} In addition to outlining the historical aspects of the principal measurements, the reviews present a critical evaluation of the possible uncertainties existing in the measurements and the analyses based on the measurements. Therefore, only a brief discussion will be presented herein concerning the principal contributions that are finding widest acceptance today. These are the contributions of Johnson⁶ and Nicolet.^{7,8}

In 1954, Johnson published an analysis of available solar measurements and a listing of extraterrestrial solar spectral irradiance at the earth's mean distance from the sun. This listing has been in wide use to the present time. The data in different spectral regions are attributable to different researchers whose instrumentation and absolute accuracy varied widely. At wavelengths longer than 600 nm, values based on Moon's analysis⁹ of previous measurements, primarily those performed by the Smithsonian Institution,¹⁰ were used. Moon's results contain the assumption that the sun radiates as a 6000-K greybody for all wavelengths beyond 1250 nm. The measurements of Dunkelman and Scolnik,¹¹ adjusted upward 9% to agree with the Smithsonian data at 600 nm and to incorporate a radiometric scale change, covered the spectral region down to 318 nm. Rocket spectrograph measurements obtained by the Naval Research Laboratory¹² were used to 220 nm. These short-wavelength data were obtained on a relative energy scale and were adjusted to the Dunkelman and Scolnik data at 318 nm. Further rocket measurements¹³ with improved instrumentation and calibration procedures have revised these early data. However, the subsequent measurements were still adjusted to the revised Dunkelman and Scolnik data in the 300–330 nm spectral region.

In 1951, Nicolet derived extraterrestrial solar spectral irradiance by noting that the intensity at the center of the disk could be closely approximated by a smooth 7200-K continuum with superimposed Fraunhofer absorption. He applied corrections for the fraction attenuated due to Fraunhofer absorption and for the variation from the center to the limb of the disk. In the region below 300 nm, he used early NRL rocket data directly.

A major problem associated with Nicolet's analysis is that the continuum cannot be directly located in regions of intense Fraunhofer absorption. This is particularly important in the region between 300 nm and 450 nm and is reflected in a greater uncertainty in this region than at longer wavelengths. A comparison of solar spectral

Table I. Flight Parameters

Flight date	Type of flight	Solar zenith angle, deg.	Altitude, km	Air mass
1967				
Aug. 3	Afternoon	56.4–72.5	11.7	0.368–0.672
Aug. 8	Sunset	51.7–80.0	11.6	0.329–1.180
Aug. 10	Sunrise	79.9–60.2	11.6	1.150–0.410
Aug. 14	Afternoon	34.3–50.0	11.6	0.247–0.318
Aug. 16	Solar transit	27.8–31.0	11.6	0.231–0.218
Aug. 19	Afternoon	28.3–43.4	11.6	0.212–0.281
Aug. 22	Solar transit	23.5–26.0	11.6	0.193–0.227
Oct. 10	Solar transit	33.5–34.3	11.6	0.243–0.247
Oct. 12	Solar transit	29.8–30.4	12.0	0.210–0.217
Oct. 22	Solar transit	22.7–23.5	12.2	0.203–0.204
Nov. 1	Solar transit	53.6–64.0	12.5	0.374–0.433

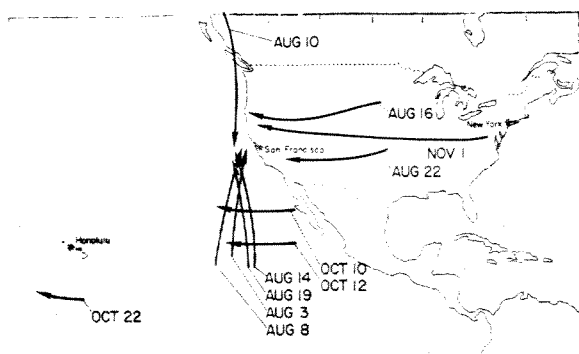


Fig. 3. Flight paths during 1967 solar observations.

irradiance as derived by Nicolet and Johnson is shown in Fig. 1.

III. Solar Spectral Irradiance Measurements

The measurements performed from the NASA CV-990 aircraft circumvented many of the difficulties encountered by previous investigators. The excellent observational conditions at flight altitude allowed the spectroradiometer system to be used to its full capabilities.

A brief description of the aircraft and the observational conditions during the experiment will be presented to illustrate the suitability of high-altitude aircraft for solar studies. This description will be followed by a discussion of the spectroradiometer system and the data reduction technique.

A. Aircraft and Observational Conditions

A specially modified four-engine jet aircraft was used for the experiment. Figure 2 is a photograph of the aircraft in flight. With an equipment payload of 7000 kg, the aircraft is capable of providing 4 h of stable observation at 11 km or 2 h at about 12 km at a speed of approximately 900 km/h. During solar observation, the aircraft heading is 90° to the solar vector and can be maintained to within $\pm \frac{1}{4}^\circ$. The stability in roll is within $\pm \frac{3}{4}^\circ$. The overall stability is well within the $\pm 3^\circ$ required for the measurements.

Flights were made at various times of day to achieve a wide variation in atmospheric path length so that the measurements could be extrapolated outside the atmosphere. Latitude and longitude of the observations were measured at 10-min intervals to an accuracy of 1 min of arc. From these data, and tabulated ephemeris data,¹⁴ a continuous determination of solar zenith angle vs time was computed. All observational data were related to Universal Time as a reference base. Information pertinent to the flights is given in Table I.

The flight paths during observational periods are shown in Fig. 3. The primary base of operations was Moffett Field, California, near San Francisco. A typical data flight profile would include a climbing outbound leg to the starting point of observations and an observation run at constant solar bearing on the

homeward leg. One exception was the final 1 November solar transit flight that originated on the east coast of the United States and terminated on the west coast.

Figure 4 shows the arrangement within the aircraft of the spectroradiometer and mirror used in the experiment. Solar observations were made from the left side of the aircraft through either an upper (65° elevation) window or a side (14° elevation) window, depending on the solar zenith angle. The windows were 2.5-cm thick fused-silica flats (Corning 7940) with clear apertures of 30 cm by 36 cm and 25 cm by 36 cm for the upper and side, respectively. Windows were cleaned prior to takeoff and externally sealed with removable shutters until observational altitude was achieved. Warm air from the cabin air conditioner was directed over the interior of the windows to prevent the condensation of cabin moisture, since the outside air temperature is about -50°C . The mirror was a 40-cm diam aluminized optical flat. An accurate calibration of the reflectance of the mirror and the transmittance of the windows as a function of wavelength and angle of incidence is extremely important for absolute intensity measurements. A detailed discussion of the calibration procedures for the windows and mirror is presented in Appendix A.

Spectral irradiance measurements were made from pressure altitudes between 11.6 km and 12.5 km during eleven data flights. Pressure altitude, rather than geometric altitude, was used in the data analysis as it is indicative of the total amount of atmosphere above the flight level. In general, the atmosphere showed excellent clarity and stability. The atmospheric transmission values derived from the data clearly bear out this point. At these altitudes, the observer is above 80% of the permanent gases in the atmosphere and virtually all variable components, including water vapor, clouds, smoke, haze, and dust. Measurements of atmospheric precipitable water vapor above the flight level indicated a total path length ranging from 8×10^{-4} precipitable-cm to 14×10^{-4} precipitable-cm.

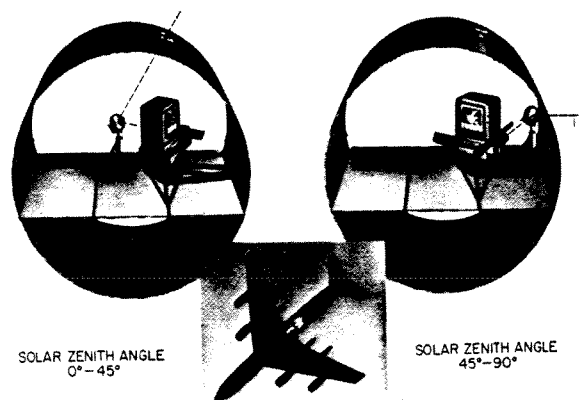


Fig. 4. Aft view showing arrangement of spectroradiometer and mirror within the aircraft for observations through the upper window ($0^\circ < z < 45^\circ$) and side window ($45^\circ < z < 90^\circ$). Inset shows spectroradiometer location within the aircraft.

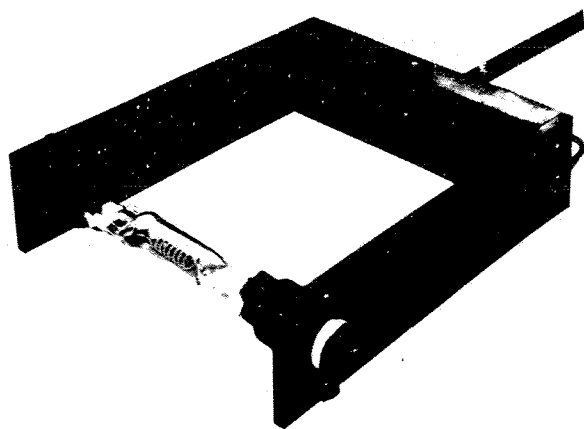


Fig. 5. Standard of spectral irradiance mounted in its holder.

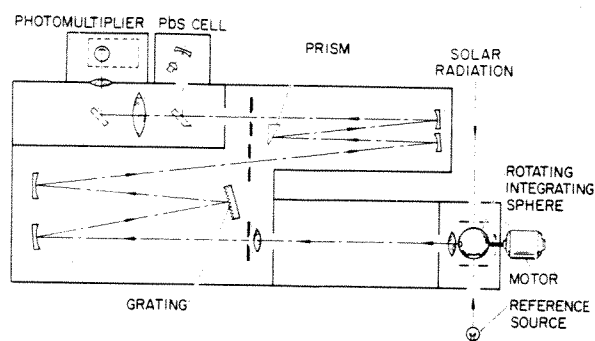


Fig. 6. Optical schematic drawing of Cary spectroradiometer showing rotating integrating sphere for alternate viewing of reference and solar radiation.

Approximately 4×10^{-4} precipitable-cm was in the path of observation within the aircraft. The corrections due to such low amounts of water vapor had negligible effect on the results. The total CO_2 content above flight altitude was approximately 40 atm-cm, while 0.5 atm-cm was attributable to CO_2 within the aircraft. Small corrections to the data were necessary at the strong absorption bands centered at 1960 nm, 2010 nm, and 2060 nm. The flights were conducted in the summer and fall when the ozone concentration is near minimum in its annual cycle. The total ozone content above an altitude of 11.6 km was derived from atmospheric transmission measurements to be 0.24 atm-cm, based on the ozone absorption coefficients of Vigroux.¹⁵

B. Spectroradiometer System

The establishment of the Standard of Spectral Irradiance in 1963 by the National Bureau of Standards² provided a common basis for spectral irradiance measurements that can be conveniently and accurately used. This Standard is a 1000-W lamp consisting of a coiled tungsten filament (3000 K) enclosed in a 1-cm by 7-cm quartz envelope containing a small amount of halogen gas. The Standard is issued at the present time through The Eppler Laboratory, Newport,

Rhode Island. Figure 5 is a photograph of the lamp in its holder. A calibration is performed over the 250–2500 nm wavelength interval at a distance of 50 cm from the lamp, based on the spectral radiance of a blackbody computed from Planck's equation.

The spectroradiometer used is a recently developed instrument whose calibration is based directly on the Standard of Spectral Irradiance. A schematic drawing of the instrument is shown in Fig. 6 and a detailed description is given by Hedelman and Nelson.¹⁶ The spectroradiometer is a modification of the Cary Model 14 ratio-recording spectrophotometer. It utilizes a prism-grating double monochromator, providing 1-nm/mm reciprocal dispersion in the uv and 3 nm/mm-reciprocal dispersion in the near ir. Resolving power is 0.1 nm in the uv and 0.3 nm in the near ir. Stray light is extremely low, less than 0.0001% throughout much of its operating range.

The most significant feature of the spectroradiometer is the incorporation of a 5-cm diam rotating integrating sphere that alternately views the reference and solar sources at a frequency of 30 Hz (see Fig. 6). The interior of the sphere is coated with magnesium oxide to diffuse the incoming radiation and provide a source of uniform intensity at the entrance to the monochromator. This feature eliminates the different geometry existing between collimated solar radiation and the divergent radiation from the reference source. This differs from the widely used practice of reimaging the source on the entrance slit, which has been determined to be a primary factor contributing to the inaccuracies previously associated with spectral irradiance measurements.³ The rotating integrating sphere, in addition to diffusing the incoming radiation, serves as an optical chopper for the solar and reference beams.

Thus, the ratio between the two sources is measured at 30 Hz, and all errors due to detector and amplifier drift are minimized. The detector-amplification system is capable of measuring the ratio of the monochromatic intensities of the reference and solar sources to an accuracy of 0.2%.

A photograph of the spectroradiometer in operation during one of the flights is shown in Fig. 7. A spectral



Fig. 7. Photograph of spectroradiometer system in operation during solar observation.

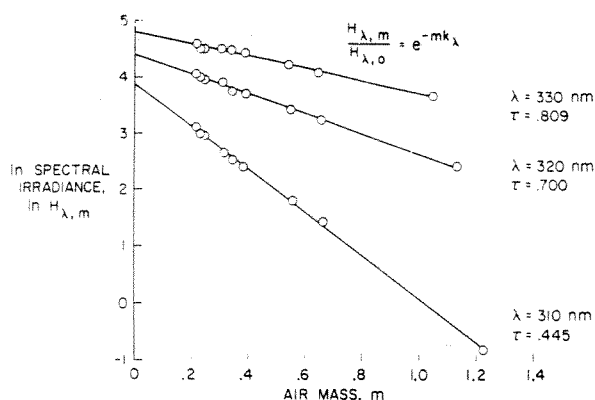


Fig. 8. Extrapolation procedure based on the Langley method.

scan took approximately 2 h to cover the wavelength range from 290 nm to 2500 nm. A 1P28 photomultiplier was used below 700 nm and a lead sulfide cell was used at longer wavelengths. In addition to a real-time strip chart display of the data, spectral irradiance, wavelength, and slitwidth were recorded every 0.1-nm wavelength interval by a Datex Model SDS-1 digital data acquisition system and a Teletype Model BRPE high-speed paper-tape punch. Data were transferred to magnetic tape on a Honeywell 200 computer and subsequently analyzed on IBM 7040-7094DCS and IBM System/360 computers.

C. Data Reduction

The classical Langley method¹⁷ was used to extrapolate the measurements to extraterrestrial conditions. This method, in addition to giving the extraterrestrial spectral irradiance level, yields atmospheric transmission coefficients from flight altitude to space.

The spectral irradiance at flight altitude, $H_{\lambda,m}$, is related to the extraterrestrial spectral irradiance level, $H_{\lambda,0}$, by

$$H_{\lambda,m} = H_{\lambda,0} e^{-mk_{\lambda}} \quad (1)$$

or

$$\ln H_{\lambda,m} = \ln H_{\lambda,0} - mk_{\lambda}. \quad (2)$$

The absolute air mass m is defined as the path length of atmosphere between the observer and the sun relative to the vertical path length at sea level. The atmospheric attenuation coefficient k_{λ} indicates the degree of absorption and scattering above the observational altitude.

The method of Kasten,¹⁸ corrected for flight altitude, was used to evaluate the variation in relative air mass as a function of solar zenith angle. Air mass varies as the secant of the solar zenith angle, z , with higher order corrections due to atmospheric curvature and refraction. As shown by Kasten, the greatest contribution to these corrections results from the lower atmosphere.

The absolute air mass used in data reduction was obtained by the relationship

$$m = (P/P_0)[\sec z + (P/P_0)(m_k - \sec z)], \quad (3)$$

where m_k is the relative air mass at sea level as derived by Kasten and P/P_0 is the ratio of the pressure at altitude to the standard sea level pressure of 760 Torr. The pressure ratio P/P_0 was obtained with high precision from one observation to another by means of the aircraft altimeter, which has resolution better than 1 part in 2000. At altitudes above 11.6 km the correction term, $P/P_0 (m_k - \sec z)$, is smaller than 0.5% of $\sec z$ for $0^\circ \leq z \leq 78^\circ$.

The spectral irradiance at observational altitude, $H_{\lambda,m}$, was directly determined from the measured spectral irradiance within the aircraft, $H^*_{\lambda,m}$, by accounting for attenuation at the window and mirror. A further correction of the data to the earth's mean distance from the sun was also made. The result can be expressed as

$$H_{\lambda,m} = (a_0^2 / \rho_{\lambda} \tau_{\lambda}) H^*_{\lambda,m}. \quad (4)$$

The values of window transmittance τ_{λ} and mirror reflectance ρ_{λ} were measured at their respective angles of incidence by methods described in Appendix A. The solar radius vector a_0 was obtained from Ref. 14 for the particular date of the measurement.

Table II. Atmospheric Transmittance from Space to 11.6 km (Zenith Direction)

Wavelength, nm	Transmittance	Wavelength, nm	Transmittance
298	0.036	650	0.960
299	0.055	700	0.972
300	0.083	757	0.976
302	0.134	761	0.913
304	0.195	768	0.980
306	0.285	800	0.980
308	0.371	900	0.988
310	0.445	1000-1930	0.99+
315	0.595	1940	0.972
320	0.700	1950	0.935
325	0.763	1960	0.920
330	0.809	1970	0.988
335	0.822	1980	0.99+
340	0.840	1990	0.99+
350	0.864	2000	0.889
360	0.879	2010	0.887
380	0.891	2020	0.982
400	0.897	2030	0.99+
450	0.925	2040	0.99+
500	0.940	2050	0.966
550	0.948	2060	0.920
575	0.932	2070	0.959
600	0.932	2080-2500	0.99+
625	0.942		

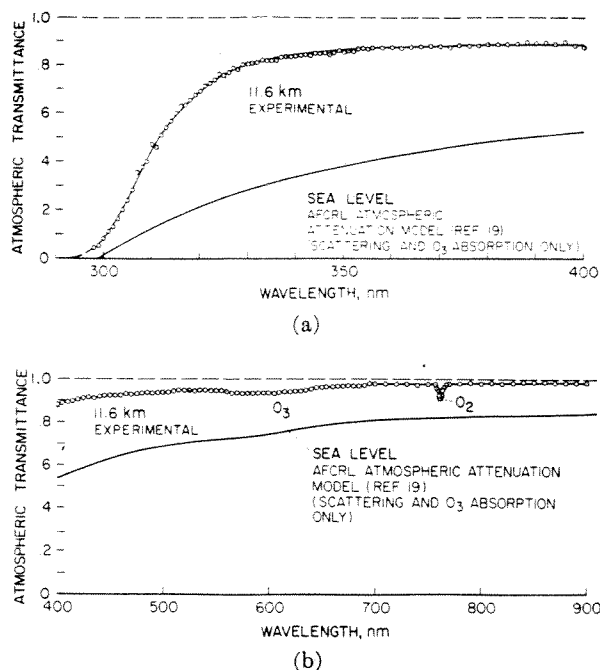


Fig. 9. Atmospheric transmittance to space from an altitude of 11.6 km: (a) uv region, (b) visible, near ir region.

Equation (2) shows the linear relationship that exists between the natural logarithm of spectral irradiance and air mass. Extrapolations to zero air mass were performed by a least-squares linear fit to the data at selected wavelengths from 300 nm to 2500 nm. Figure 8 illustrates the extrapolation procedure used. Data from each flight at several wavelengths in the Hartley-Huggins ozone bands are plotted to show the rapidly decreasing transmittance in the uv due to ozone absorption in the upper atmosphere. The variability of the ozone from flight to flight, although small, results in significantly more scatter in the data below 340 nm than at longer wavelengths. The ordinate intercept yields the natural logarithm of extraterrestrial solar spectral irradiance at the earth's mean distance from the sun, and the negative of the slope yields the atmospheric attenuation coefficient, k_λ .

Vertical atmospheric transmittance values from space to an altitude of 11.6 km ($m = 0.204$) were derived from the measured values of k_λ by the relationship

$$\tau_{\lambda,m} = e^{-mk_\lambda}. \quad (5)$$

These values are tabulated in Table II and plotted in Figs. 9(a) and 9(b) for the uv and visible-near ir regions, respectively.

Atmospheric transmittance was found to be above 0.99 for essentially all wavelengths longer than 1000 nm and greater than 0.90 for wavelengths longer than 410 nm. The average deviation of the experimental atmospheric transmittance values from a smoothed curve varies from 0.02 in the ozone absorption region below 340 nm to less than 0.005 at longer wavelengths. The Chappuis ozone band between 550 nm and 600 nm,

resulting in a decrease in atmospheric transmittance of less than 2%, is readily seen in the measurements. The only atmospheric absorption detected, other than ozone, was O₂ at 760 nm and CO₂ at 1960 nm, 2010 nm, and 2060 nm. Water vapor absorption at 1380 nm was not detected due to the presence of absorption by the fused-silica window within the same wavelength region. As shown in Fig. 9(a), ozone absorption becomes apparent for wavelengths shorter than 340 nm and virtually complete at wavelengths shorter than 295 nm. The practical limit of meaningful measurements is 300 nm.

The advantage of performing measurements from an aircraft or other high-altitude platform is obvious when a comparison is made with the transmittance to sea level. Such a comparison is made in Figs. 9(a) and 9(b) between the experimental atmospheric transmittance values from 11.6 km altitude to space and the values from sea level to space predicted by the AFCL Atmospheric Attenuation Model.¹⁹ For example, at 340 nm the transmittance to space from 11.6 km is 0.84 compared with a maximum of 0.33 from sea level. The high atmospheric transparency at all wavelengths beyond the primary ozone absorption region enables extrapolation to zero air mass to be performed with a high degree of accuracy.

IV. Experimental Results

The resultant extraterrestrial spectral irradiance in the 300–2500 nm wavelength interval, obtained from extrapolation of all measurements to zero air mass, is presented in both tabular and graphical form in Appendix B. The wavelength interval in Table VII, which is somewhat representative of the spectral resolution of the data, varies from 0.1 nm over most of the uv and visible to 5 nm in the smooth continuum region of the ir. This relatively high resolution allows identification of the major Fraunhofer lines through the use of the Rowland Table.²⁰ The identities of selected major absorption lines are indicated in the spectrum.

Most applications of solar spectral irradiance do not require resolution as high as that presented in Appendix B. For that reason, the percentage of total irradiance at wavelengths shorter than the printout wavelength is included in the listing to provide a simple method of calculating the total irradiance or average spectral irradiance within an arbitrary wavelength interval.

A listing of extraterrestrial solar spectral irradiance, smoothed over relatively wide wavelength intervals, will be presented for engineering use and the uncertainties in the spectral irradiance values will be discussed in the following sections.

A. Solar Spectral Irradiance Listing for Engineering Use

The engineering technology presently being applied to space research does not, in general, require the relatively high resolution given in Appendix B. In many instances it is desirable to use values of solar spectral

Table III. Extraterrestrial Solar Spectral Irradiance at the Earth's Mean Distance from the Sun

λ , nm	H_{λ} , W/m ² -nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² -nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² -nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² -nm	P_{λ} , %
200*	0.012	0.01	375	1.13	8.12	600	1.79	35.69	950	0.799	67.06
205*	0.016	0.02	380	1.11	8.56	610	1.78	36.98	960	0.792	67.63
210*	0.026	0.03	385	1.05	8.92	620	1.72	38.23	970	0.781	68.29
215*	0.043	0.04	390	1.10	7.32	630	1.70	39.47	980	0.776	68.76
220*	0.055	0.06	395	1.21	7.71	640	1.67	40.68	990	0.758	69.31
225*	0.062	0.08	400	1.34	8.19	650	1.61	41.85	1000	0.741	69.85
230*	0.064	0.10	405	1.77	8.82	660	1.56	42.98	1050	0.669	72.39
235*	0.057	0.12	410	1.83	9.46	670	1.58	44.13	1100	0.602	74.66
240*	0.061	0.14	415	1.88	10.14	680	1.57	45.27	1150	0.546	76.72
245*	0.069	0.17	420	1.85	10.82	690	1.49	46.36	1200	0.496	78.59
250*	0.064	0.19	425	1.76	11.47	700	1.50	47.45	1250	0.451	80.30
255*	0.100	0.22	430	1.67	12.08	710	1.43	48.50	1300	0.419	81.96
260*	0.125	0.26	435	1.75	12.67	720	1.39	49.51	1350	0.379	83.39
265*	0.178	0.31	440	1.92	13.34	730	1.37	50.51	1400	0.349	84.60
270*	0.22	0.39	445	2.01	14.05	740	1.32	51.48	1450	0.329	85.81
275*	0.20	0.46	450	2.06	14.79	750	1.30	52.42	1500	0.296	86.91
280*	0.21	0.54	455	2.10	15.54	760	1.28	53.35	1600	0.253	88.88
285*	0.30	0.63	460	2.11	16.30	770	1.22	54.25	1700	0.215	90.37
290*	0.46	0.76	465	2.03	17.06	780	1.21	55.12	1800	0.170	91.85
295*	0.56	0.95	470	2.07	17.80	790	1.19	55.99	1900	0.139	93.03
300	0.56	1.15	475	2.11	18.55	800	1.16	56.83	2000	0.119	93.97
305	0.59	1.35	480	2.10	19.31	810	1.12	57.64	2100	0.0945	94.74
310	0.66	1.57	485	1.98	20.06	820	1.08	58.43	2200	0.0764	95.34
315	0.73	1.83	490	1.96	20.74	830	1.06	59.20	2300	0.0644	95.56
320	0.75	2.10	495	2.00	21.47	840	1.02	59.95	2400	0.0575	96.29
325	0.89	2.37	500	1.96	22.18	850	0.966	60.67	2500	0.0532	96.66
330	1.04	2.74	510	1.97	23.60	860	0.961	61.36	3000*	0.0266	97.96
335	1.02	3.12	520	1.85	24.94	870	0.942	62.04	3500*	0.0139	98.67
340	1.09	3.47	530	1.92	26.31	880	0.936	62.72	4000*	0.0094	99.05
345	1.01	3.84	540	1.89	27.69	890	0.917	63.38	4500*	0.0061	99.32
350	1.04	4.19	550	1.88	29.05	900	0.893	64.03	5000*	0.0041	99.49
355	1.02	4.58	560	1.83	30.33	910	0.866	64.66	7000*	0.0011	99.81
360	0.98	4.93	570	1.85	31.70	920	0.842	65.28	10000*	0.0003	99.92
365	1.15	5.29	580	1.88	33.05	930	0.840	65.88	20000*	0.0002	99.98
370	1.15	5.75	590	1.83	34.38	940	0.818	66.48	40000*	0.0001	99.99

*Adjusted NRL data (Ref. 13)
 †5500°K greybody

irradiance that have been smoothed over the detailed Fraunhofer structure. Such a listing is provided in Table III as a solar spectral irradiance listing proposed for engineering use.

The values of spectral irradiance in Table III corresponding to wavelengths shorter than 1000 nm have been integrated over a 10-nm interval centered at the respective wavelength. At longer wavelengths, the data have been integrated over the printout interval. Integration was performed numerically at 0.1-nm increments. A listing of the percentage P_{λ} of the total solar irradiance at wavelengths shorter than the printout wavelength is also provided.

In the 300–2500 nm wavelength region, the results are from the present experiment. Below 300 nm, the most recent NRL rocket spectrograph measurements, reported by Tousey,¹³ are used. These data, obtained originally on a relative scale, have been normalized to the present experiment in the overlap region 300–400 nm. Beyond 2500 nm, a 5800-K greybody approximation is believed to be an adequate representation of the solar continuum.

The total irradiance, calculated from the integral of the spectral irradiance, is 1390 W/m² (1.99 cal/cm²-min). This number represents the total solar irradiance incident on a surface in space oriented normal to the sun at the earth's mean distance.

B. Uncertainty Limits of the Results

The measurements presented in this paper are based on the standard of spectral irradiance maintained by the National Bureau of Standards. The maximum uncertainty existing in the standard varies from approximately 5% at 300 nm to 3% in the visible and ir.² Total irradiance could possibly be more accurate since integration over a number of spectral data points tends to cancel the random component of error. The degree of systematic error is not known, although, as discussed later, results indicate that it probably does not exceed 2%.

The precision of the measurements is better than 1%. The absolute accuracy of the resultant extraterrestrial solar spectral irradiance can be determined from the individual accuracies of each factor affecting the measurement. The uncertainty in the measurement of a quantity $\xi(X)$ whose value is a function of n independent variables $X_1, X_2, \dots, X_n$, each with an uncertainty $\Delta X_1, \Delta X_2, \dots, \Delta X_n$, can be obtained by

$$\Delta \xi = \left[\left(\frac{\partial \xi}{\partial X_1} \Delta X_1 \right)^2 + \left(\frac{\partial \xi}{\partial X_2} \Delta X_2 \right)^2 + \dots + \left(\frac{\partial \xi}{\partial X_n} \Delta X_n \right)^2 \right]^{1/2} \quad (6)$$

Equations (1), (4), and (5) may be combined to express the extraterrestrial spectral irradiance $H_{\lambda,0}$ in terms of

Table IV. Uncertainty Limits

λ , nm	Standard of Spectral Irradiance, %	Spectroradiometer measurement, %	Zero air mass extrapolation, %	Mirror reflectance, %	Window transmittance, %	Total uncertainty, %
300	5.0	0.8	24.0	0.5	0.5	25.0
310	5.0	0.7	5.0	0.5	0.4	7.1
320	5.0	0.7	3.0	0.5	0.4	5.9
330	4.0	0.7	2.5	0.5	0.4	4.8
350	4.0	0.6	1.2	0.5	0.4	4.3
400	3.0	0.5	0.6	0.5	0.4	3.2
500	3.0	0.5	0.5	0.5	0.4	3.1
700	3.0	0.5	0.5	0.5	0.4	3.1
1000	3.0	0.5	0.3	0.5	0.4	3.1
1500	3.0	0.7	0.2	0.5	0.4	3.2
2000	3.0	0.8	0.2	0.5	0.4	3.2
2200	3.0	1.5	0.2	0.5	1.6	3.8
2500	3.0	2.0	0.2	0.5	0.9	3.8

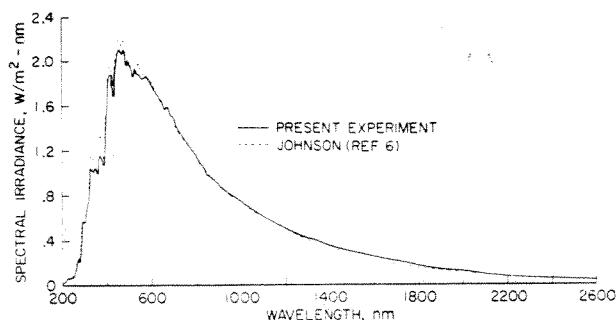


Fig. 10. Comparison of solar spectral irradiance, present experiment and Johnson.

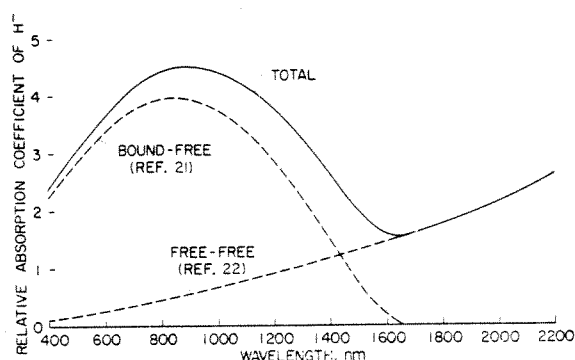


Fig. 11. Spectral variation of absorption coefficient of the negative hydrogen ion (H^-).

the measured spectral irradiance $H_{\lambda,m}^*$, atmospheric transmittance $\tau_{\lambda,m}$, distance from the sun a_0 , mirror reflectance ρ_λ , and window transmittance τ_λ :

$$H_{\lambda,0} = \frac{a_0^2 H_{\lambda,m}^*}{\tau_{\lambda,m} \rho_\lambda \tau_\lambda} \quad (7)$$

The fractional uncertainty in $H_{\lambda,0}$ can now be expressed in terms of the fractional uncertainties in the individual factors:

$$\frac{\Delta H_{\lambda,0}}{H_{\lambda,0}} = \left[\left(\frac{\Delta H_{\lambda,m}^*}{H_{\lambda,m}^*} \right)^2 + \left(\frac{\Delta \tau_{\lambda,m}}{\tau_{\lambda,m}} \right)^2 + \left(\frac{2\Delta a_0}{a_0} \right)^2 + \left(\frac{\Delta \rho_\lambda}{\rho_\lambda} \right)^2 + \left(\frac{\Delta \tau_\lambda}{\tau_\lambda} \right)^2 \right]^{\frac{1}{2}} \quad (8)$$

Table IV lists the individual uncertainties and the total uncertainty in $H_{\lambda,0}$ at various wavelengths from 300 nm to 2500 nm. The uncertainty associated with distance from the sun is small enough to be neglected with respect to the others and was not included in the table.

The largest total uncertainty occurs in the region of high ozone absorption below 320 nm. The uncertainty quickly decreases at longer wavelengths as the atmospheric transmittance increases and remains approximately 3% throughout the entire visible and ir. The primary uncertainty at wavelengths longer than 320

nm is due to uncertainty in the Standard of Spectral Irradiance.

V. Comparison With Previous Studies

A comparison between the values of spectral irradiance listed in Table III and Johnson's values⁶ is shown in Fig. 10. In general, the results are relatively close considering the diverse data used by Johnson in his analysis. The primary region of discrepancy is in the uv where the results of the present experiment indicate a uniform 11% lower irradiance between 300 nm and 400 nm. There are regions of slight discrepancy in the visible and near ir and a broad region between 1300 nm and 1900 nm where the continuum is significantly higher than Johnson's level.

In the ir at wavelengths beyond 1250 nm, Johnson accepted the assumption that the sun radiated as a 6000-K greybody. This means that the spectral absorptance of the solar atmosphere is assumed to be independent of wavelength. This is not the case, and, in fact, as shown in Fig. 11, the absorptance of the solar atmosphere has a minimum in the same ir region as the results from the present experiment indicate higher irradiance.*

The lower intensity level in the spectral region below 600 nm is in good agreement with recently published

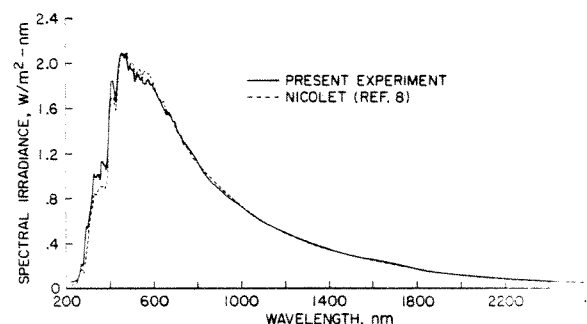


Fig. 12. Comparison of solar spectral irradiance, present experiment and Nicolet.

* The data in Fig. 11 were obtained from Refs. 21 and 22 and represent known spectral variations in the opacity of the solar atmosphere. Most of the continuum absorption is due to the negative ion of hydrogen (H^-), which absorbs energy through either "bound-free" transitions or "free-free" transitions which are functions of wavelength. The minimum in the total absorption coefficient between 1300 nm and 1900 nm means that the radiation is emanating from a higher temperature region deeper within the solar atmosphere. Beyond 2500 nm, the free-free absorption coefficient continues to rise, with the result that the effective greybody temperature of the sun decreases at longer wavelengths. The exact relationship between the absorption coefficient and the effective solar temperature in the ir is not known, however, since the effect is small, the assumption made in the present experiment that a 5800-K solar temperature is an adequate representation at wavelengths longer than 2500 nm.

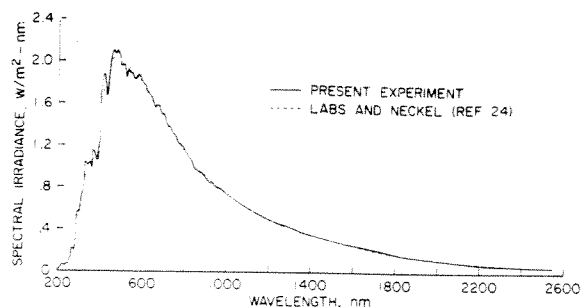


Fig. 13. Comparison of solar spectral irradiance, present experiment and Labs and Neckel.

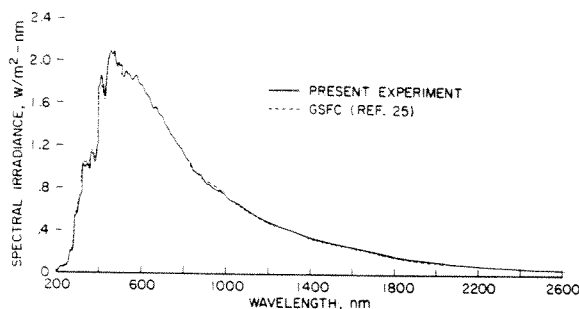


Fig. 14. Comparison of solar spectral irradiance, present experiment and Goddard Space Flight Center (GSFC).

results of measurements between 310 nm and 530 nm by Stair and Ellis obtained at the Mauna Loa Observatory, Hawaii.²³ Using Johnson's values of spectral irradiance at wavelengths longer than 530 nm, they concluded that the total irradiance is approximately 2% less than Johnson's value of 1395 W/m². However, the results of the present experiment indicate that the higher level of spectral irradiance in the ir would nearly compensate for the lower level in the uv and visible, with the result that the value of total irradiance should be virtually the same as Johnson's value.

The spectral irradiance values of Nicolet are compared with those of the present experiment in Fig. 12. There is general agreement in the visible and ir, except in the 500–600 nm region, where Nicolet's values appear to be an average of 3% higher than the present results. The primary region of discrepancy is in the uv. The difficulty of determining the true level of Fraunhofer absorption in this region, a necessary factor in Nicolet's method, is felt to be the probable cause of the discrepancy.

The results of Labs and Neckel,²⁴ obtained at the Jungfraujoch Scientific Station, Switzerland (altitude 3.6 km), are shown in Fig. 13. Their data are the results of measurements of the spectral radiance at the center of the solar disk in the 330–1250 nm wavelength region adjusted for center-limb variations and line blanketing to obtain mean spectral radiance. In the uv below 330 nm, they adjusted the data given by Tousey¹³ and, in the ir, used the normalized continuum

from solar models with corrections for the ratio of mean to central intensity and line blanketing. The agreement with results from the present experiment is very close except in the region below 500 nm where the results of Labs and Neckel average about 5% lower. Possible explanations for this discrepancy include inaccuracies in the evaluation of the higher degree of atmospheric attenuation experienced in their observations and inaccuracies in the determination of the center-limb variations and line blanketing coefficients. The value of total irradiance obtained by Labs and Neckel is 1366 W/m².

Researchers from the NASA Goddard Space Flight Center (GSFC), participating with independent instrumentation during seven of the CV-990 flights, have published values²⁵ of extraterrestrial solar spectral irradiance obtained from weighted averages of measurements from four instruments: (1) Perkin-Elmer model 112 monochromator; (2) modified Eppley filter radiometer; (3) Carl Leiss monochromator, and (4) Block model P-4 polarization interferometer. The measurements obtained with the latter two instruments were on a relative basis due to calibration difficulties. A comparison between the GSFC results and the results from the present experiment is presented in Fig. 14. The agreement is close, except in the spectral region between 500 nm and 800 nm, where the GSFC results indicate considerably lower spectral irradiance. The reason for this discrepancy is not clear at the present time, although, as shown in Ref. 25, the differences between the individual GSFC instruments in this spectral region are of this magnitude. The value of total irradiance obtained by GSFC is 1351 W/m².

The value of total irradiance obtained from the present experiment, 1390 W/m², can be compared to a recently published value of 1360 W/m² based on the International Pyrheliometric Scale. This value is the result of a total radiometer experiment by Jet Propulsion Laboratory and Eppley Laboratory aboard the X-15 aircraft.²⁶ The relatively close agreement between the two total irradiance values, which are based on completely different radiometric standards, indicates

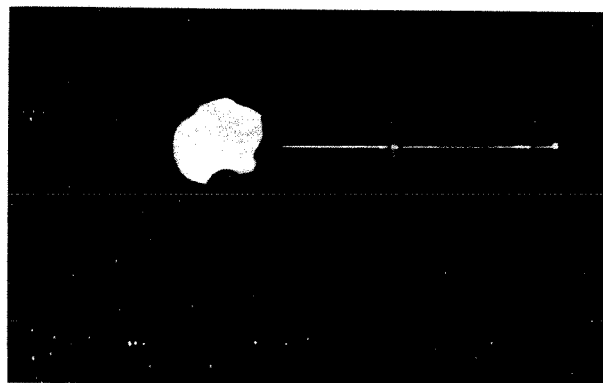


Fig. 15. Apparatus for measuring transmittance of fused silica windows.

Table V. Transmittance of 2.5-cm Fused Silica as a Function of Wavelength and Angle of Incidence (Unpolarized Radiation)

Wavelength, nm	Angle of incidence			Wavelength, nm	Angle of incidence		
	0°	20°	30°		0°	20°	30°
290	0.860	0.856	0.850	1410	0.873	0.869	0.859
295	0.864	0.860	0.855	1420	0.898	0.896	0.890
300	0.877	0.873	0.869	1430	0.916	0.915	0.912
320	0.910	0.909	0.906	1500	0.942	0.941	0.940
350	0.926	0.925	0.923	1600	0.943	0.943	0.943
400	0.937	0.937	0.937	1700	0.940	0.940	0.940
450	0.939	0.939	0.939	1800	0.936	0.936	0.935
500	0.939	0.939	0.939	1900	0.937	0.937	0.937
550	0.940	0.940	0.940	2000	0.937	0.937	0.937
600	0.941	0.941	0.941	2100	0.929	0.929	0.927
700	0.942	0.942	0.942	2140	0.898	0.895	0.889
800	0.942	0.942	0.942	2160	0.833	0.826	0.809
900	0.942	0.942	0.942	2180	0.753	0.744	0.715
1000	0.943	0.943	0.943	2200	0.398	0.362	0.275
1100	0.944	0.944	0.944	2220	0.292	0.249	0.145
1200	0.944	0.944	0.944	2250	0.423	0.389	0.306
1300	0.938	0.938	0.938	2270	0.535	0.508	0.443
1340	0.925	0.924	0.923	2300	0.710	0.695	0.658
1350	0.920	0.919	0.917	2350	0.827	0.820	0.802
1360	0.912	0.911	0.907	2400	0.828	0.821	0.803
1370	0.862	0.857	0.845	2420	0.781	0.771	0.745
1380	0.740	0.727	0.696	2450	0.714	0.699	0.663
1383	0.705	0.690	0.653	2470	0.680	0.663	0.621
1390	0.730	0.716	0.683	2480	0.593	0.570	0.514
1400	0.816	0.808	0.789	2500	0.514	0.486	0.417

that the systematic error in the values of extraterrestrial solar spectral irradiance is probably no greater than 2%.

The actual degree of systematic error in the measurements will be resolved when a measurement of total solar irradiance is performed in space by an absolute radiometer such as described in Ref. 27. This radiometer, the JPL Standard Total-Radiation Absolute Radiometer, is capable of better than 1.0% absolute accuracy. A direct total irradiance measurement obtained in space by such an instrument would justify a revision of the spectral irradiance data contained in this report if a significant discrepancy is found.

VI. Concluding Remarks

The NASA CV-990 aircraft afforded an excellent high-altitude platform from which to perform measurements of absolute solar spectral irradiance. Observational conditions such as stability of the aircraft and constancy of the atmosphere allowed measurements to be made with good resolution and high precision. At altitudes greater than 11.6 km, the aircraft is above 80% of the atmosphere and virtually all dust and water vapor. Therefore, the amount of atmospheric scattering was greatly reduced and the variability in atmospheric transmittance was virtually eliminated.

The primary purpose of the experiment was to reduce the uncertainties in the present knowledge of extraterrestrial solar spectral irradiance at the earth's mean distance from the sun. The measurements covered a spectral region encompassing over 95% of the total energy emitted by the sun, using a single instrument of

high precision. The results of the measurements are directly traceable to the NBS Standard of Spectral Irradiance and possess an absolute accuracy, over most of the wavelength range, of approximately 3% and resolution to 0.1 nm. These measurements, integrated over relatively wide spectral intervals, form the basis for a solar spectral irradiance table (Table III) proposed for engineering use.

The solar spectral irradiance data in Table VII, covering the 300–2500 nm spectral interval with resolution to 0.1 nm, are available to any organization that has need for such information. The data may be obtained as a listing on either cards or tape from National Space Science Data Center, Code 601, Goddard Space Flight Center, Greenbelt, Maryland 20771.

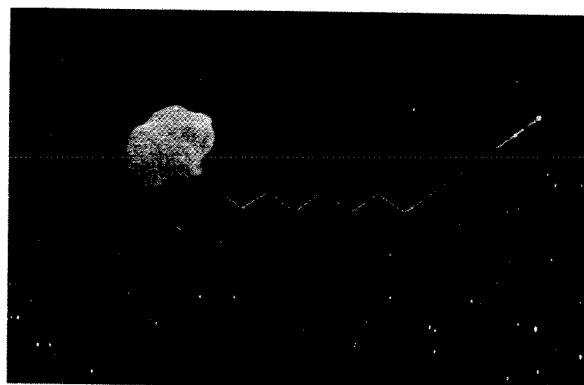


Fig. 16. Apparatus for measuring mirror reflectance.

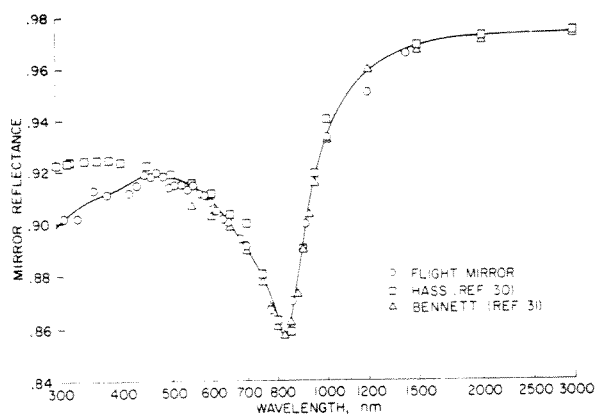


Fig. 17. Spectral reflectance of flight mirror.

Table VI. Reflectance of Front-Surface Aluminum Mirror as a Function of Wavelength and Angle of Incidence (Unpolarized Radiation)

Wavelength, nm	Angle of incidence			
	0°	35°	45°	55°
300	0.898	0.897	0.896	0.894
350	0.909	0.907	0.904	0.898
400	0.914	0.912	0.910	0.904
450	0.918	0.916	0.914	0.909
500	0.917	0.916	0.914	0.908
550	0.915	0.915	0.913	0.907
600	0.905	0.904	0.901	0.894
650	0.900	0.898	0.895	0.889
700	0.891	0.889	0.886	0.878
750	0.879	0.877	0.873	0.865
800	0.863	0.861	0.857	0.848
825	0.856	0.854	0.850	0.841
850	0.867	0.865	0.861	0.852
900	0.892	0.890	0.886	0.878
950	0.918	0.917	0.914	0.907
1000	0.933	0.932	0.930	0.926
1100	0.948	0.947	0.945	0.941
1200	0.957	0.956	0.954	0.950
1500	0.968	0.967	0.966	0.963
2000	0.972	0.971	0.970	0.967
2500	0.974	0.973	0.972	0.970

Appendix A. Determination of Transmittance of Windows and Reflectance of Mirror

An accurate determination of the transmittance of the fused silica windows and the reflectance of the mirror used in the measurements is highly important since an error in these values is a direct error in the measured value of solar spectral irradiance. Direct solar radiation at 11.6 km was found to be natural or unpolarized within detection limits of $\pm 0.2\%$. This determination was made by rotating a polarizer while the solar vector was at normal incidence to the upper aircraft window. At the maximum solar incidence angle of 31° , the degree of polarization of the solar radiation transmitted through the window was 2% , as expected by Fresnel reflection

considerations. Reflectance from the aluminized turning mirror for unpolarized radiation would thus be in error only 0.15% at its maximum incidence angle of 55° . It is, therefore, sufficient to use transmittance and reflectance coefficients for unpolarized radiation in analysis of the data.

A. Measurement of Window Transmittance

The spectral transmittance of the fused silica window was determined by two independent procedures. The first utilized a Carl Leiss double-prism monochromator in conjunction with a 20-cm diam magnesium oxide integrating sphere and a Brower model 129 phase-lock amplifier. A collimated beam of light was directed through an aperture in the sphere (see Fig. 15) and measurements were made with the window in and out of the light beam at each wavelength setting of the monochromator. Since the entire beam always entered the sphere, it was not necessary to correct for the differences in optical path length with the window inserted or removed. The lamp exhibited a 12% polarization, which resulted in a negligible error (0.1% for 30° incidence) in the measured transmittance.

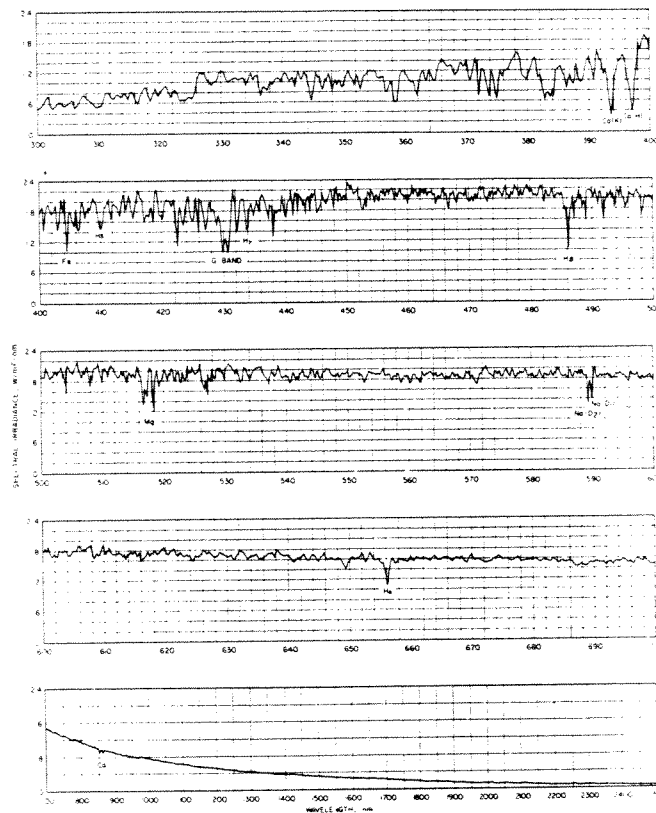


Fig. 18. Extraterrestrial solar spectral irradiance at the earth's mean distance from the sun (Table VII).

Table VII. Extraterrestrial Solar Spectral Irradiance at the Earth's Mean Distance¹ from the Sun (300-2500 nm)

λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %
300.0	0.545	1.15	340.0	1.089	3.47	380.0	1.160	4.93	420.0	1.108	6.56	460.0	1.341	7.32	500.0	1.692	8.19
302.4	0.504	1.16	342.0	1.185	3.48	382.0	1.109	4.94	402.0	1.166	6.57	442.0	1.293	7.33	482.0	1.793	8.20
304.8	0.546	1.18	344.0	1.175	3.50	384.0	1.044	4.95	404.0	1.263	7.34	444.0	1.844	7.34	484.0	1.844	8.22
307.2	0.616	1.19	346.0	1.147	3.52	386.0	0.979	4.97	406.0	1.263	7.34	446.0	1.844	7.34	486.0	1.844	8.22
309.6	0.749	1.21	348.0	1.078	3.53	388.0	0.960	4.99	408.0	1.233	7.34	448.0	1.892	7.35	488.0	1.892	8.23
312.0	0.576	1.23	350.0	1.013	3.55	390.0	0.971	5.00	410.0	1.110	7.35	450.0	1.939	7.35	490.0	1.939	8.24
314.4	0.485	1.25	352.0	1.028	3.56	392.0	0.981	5.01	412.0	1.031	7.37	452.0	1.981	7.37	492.0	1.981	8.26
316.8	0.541	1.26	354.0	1.084	3.58	394.0	0.941	5.03	414.0	1.009	7.38	454.0	1.981	7.38	494.0	1.981	8.27
319.2	0.590	1.28	356.0	1.072	3.59	396.0	0.799	5.04	416.0	1.426	7.38	456.0	1.981	7.38	496.0	1.981	8.28
321.6	0.635	1.30	358.0	0.899	3.61	398.0	0.799	5.05	418.0	1.157	7.38	458.0	1.981	7.38	498.0	1.981	8.29
324.0			360.0	0.899	3.61	400.0	1.426	6.63	420.0	1.104	7.39	460.0	1.981	7.39	500.0	1.981	8.30
326.4			362.0	0.778	3.62	402.0	1.104	6.64	422.0	1.429	7.40	462.0	1.981	7.40	502.0	1.981	8.31
328.8			364.0	0.778	3.62	404.0	1.104	6.65	424.0	1.429	7.41	464.0	1.981	7.41	504.0	1.981	8.31
331.2			366.0	0.778	3.62	406.0	1.104	6.66	426.0	1.429	7.42	466.0	1.981	7.42	506.0	1.981	8.32
333.6			368.0	0.778	3.62	408.0	1.104	6.67	428.0	1.429	7.43	468.0	1.981	7.43	508.0	1.981	8.33
336.0			370.0	0.778	3.62	410.0	1.104	6.68	430.0	1.429	7.44	470.0	1.981	7.44	510.0	1.981	8.34
338.4			372.0	0.778	3.62	412.0	1.104	6.69	432.0	1.429	7.45	472.0	1.981	7.45	512.0	1.981	8.35
340.8			374.0	0.778	3.62	414.0	1.104	6.70	434.0	1.429	7.46	474.0	1.981	7.46	514.0	1.981	8.36
343.2			376.0	0.778	3.62	416.0	1.104	6.71	436.0	1.429	7.47	476.0	1.981	7.47	516.0	1.981	8.37
345.6			378.0	0.778	3.62	418.0	1.104	6.72	438.0	1.429	7.48	478.0	1.981	7.48	518.0	1.981	8.38
348.0			380.0	0.778	3.62	420.0	1.104	6.73	440.0	1.429	7.49	480.0	1.981	7.49	520.0	1.981	8.39
350.4			382.0	0.778	3.62	422.0	1.104	6.74	442.0	1.429	7.50	482.0	1.981	7.50	522.0	1.981	8.40
352.8			384.0	0.778	3.62	424.0	1.104	6.75	444.0	1.429	7.51	484.0	1.981	7.51	524.0	1.981	8.41
355.2			386.0	0.778	3.62	426.0	1.104	6.76	446.0	1.429	7.52	486.0	1.981	7.52	526.0	1.981	8.42
357.6			388.0	0.778	3.62	428.0	1.104	6.77	448.0	1.429	7.53	488.0	1.981	7.53	528.0	1.981	8.43
360.0			390.0	0.778	3.62	430.0	1.104	6.78	450.0	1.429	7.54	490.0	1.981	7.54	530.0	1.981	8.44
362.4			392.0	0.778	3.62	432.0	1.104	6.79	452.0	1.429	7.55	492.0	1.981	7.55	532.0	1.981	8.45
364.8			394.0	0.778	3.62	434.0	1.104	6.80	454.0	1.429	7.56	494.0	1.981	7.56	534.0	1.981	8.46
367.2			396.0	0.778	3.62	436.0	1.104	6.81	456.0	1.429	7.57	496.0	1.981	7.57	536.0	1.981	8.47
369.6			398.0	0.778	3.62	438.0	1.104	6.82	458.0	1.429	7.58	498.0	1.981	7.58	538.0	1.981	8.48
372.0			400.0	0.778	3.62	440.0	1.104	6.83	460.0	1.429	7.59	500.0	1.981	7.59	540.0	1.981	8.49
374.4			402.0	0.778	3.62	442.0	1.104	6.84	462.0	1.429	7.60	502.0	1.981	7.60	542.0	1.981	8.50
376.8			404.0	0.778	3.62	444.0	1.104	6.85	464.0	1.429	7.61	504.0	1.981	7.61	544.0	1.981	8.51
379.2			406.0	0.778	3.62	446.0	1.104	6.86	466.0	1.429	7.62	506.0	1.981	7.62	546.0	1.981	8.52
381.6			408.0	0.778	3.62	448.0	1.104	6.87	468.0	1.429	7.63	508.0	1.981	7.63	548.0	1.981	8.53
384.0			410.0	0.778	3.62	450.0	1.104	6.88	470.0	1.429	7.64	510.0	1.981	7.64	550.0	1.981	8.54
386.4			412.0	0.778	3.62	452.0	1.104	6.89	472.0	1.429	7.65	512.0	1.981	7.65	552.0	1.981	8.55
388.8			414.0	0.778	3.62	454.0	1.104	6.90	474.0	1.429	7.66	514.0	1.981	7.66	554.0	1.981	8.56
391.2			416.0	0.778	3.62	456.0	1.104	6.91	476.0	1.429	7.67	516.0	1.981	7.67	556.0	1.981	8.57
393.6			418.0	0.778	3.62	458.0	1.104	6.92	478.0	1.429	7.68	518.0	1.981	7.68	558.0	1.981	8.58
396.0			420.0	0.778	3.62	460.0	1.104	6.93	480.0	1.429	7.69	520.0	1.981	7.69	560.0	1.981	8.59
398.4			422.0	0.778	3.62	462.0	1.104	6.94	482.0	1.429	7.70	522.0	1.981	7.70	562.0	1.981	8.60
400.8			424.0	0.778	3.62	464.0	1.104	6.95	484.0	1.429	7.71	524.0	1.981	7.71	564.0	1.981	8.61
403.2			426.0	0.778	3.62	466.0	1.104	6.96	486.0	1.429	7.72	526.0	1.981	7.72	566.0	1.981	8.62
405.6			428.0	0.778	3.62	468.0	1.104	6.97	488.0	1.429	7.73	528.0	1.981	7.73	568.0	1.981	8.63
408.0			430.0	0.778	3.62	470.0	1.104	6.98	490.0	1.429	7.74	530.0	1.981	7.74	570.0	1.981	8.64
410.4			432.0	0.778	3.62	472.0	1.104	6.99	492.0	1.429	7.75	532.0	1.981	7.75	572.0	1.981	8.65
412.8			434.0	0.778	3.62	474.0	1.104	7.00	494.0	1.429	7.76	534.0	1.981	7.76	574.0	1.981	8.66
415.2			436.0	0.778	3.62	476.0	1.104	7.01	496.0	1.429	7.77	536.0	1.981	7.77	576.0	1.981	8.67
417.6			438.0	0.778	3.62	478.0	1.104	7.02	498.0	1.429	7.78	538.0	1.981	7.78	578.0	1.981	8.68
420.0			440.0	0.778	3.62	480.0	1.104	7.03	500.0	1.429	7.79	540.0	1.981	7.79	580.0	1.981	8.69
422.4			442.0	0.778	3.62	482.0	1.104	7.04	502.0	1.429	7.80	542.0	1.981	7.80	582.0	1.981	8.70
424.8			444.0	0.778	3.62	484.0	1.104	7.05	504.0	1.429	7.81	544.0	1.981	7.81	584.0	1.981	8.71
427.2			446.0	0.778	3.62	486.0	1.104	7.06	506.0	1.429	7.82	546.0	1.981	7.82	586.0	1.981	8.72
429.6			448.0	0.778	3.62	488.0	1.104	7.07	508.0	1.429	7.83	548.0	1.981	7.83	588.0	1.981	8.73
432.0			450.0	0.778	3.62	490.0	1.104	7.08	510.0	1.429	7.84	550.0	1.981	7.84	590.0	1.981	8.74
434.4			452.0	0.778	3.62	492.0	1.104	7.09	512.0	1.429	7.85	552.0	1.981	7.85	592.0	1.981	8.75
436.8			454.0	0.778	3.62	494.0	1.104	7.10	514.0	1.429	7.86	554.0	1.981	7.86	594.0	1.981	8.76
439.2			456.0	0.778	3.62	496.0	1.104	7.11	516.0	1.429	7.87	556.0	1.981	7.87	596.0	1.981	8.77
441.6			458.0	0.778	3.62	498.0	1.104	7.12	518.0	1.429	7.88	558.0	1.981	7.88	598.0	1.981	8.78
444.0			460.0	0.778	3.62	500.0	1.104	7.13	520.0	1.429	7.89	560.0	1.981	7.89	600.0	1.981	8.79
446.4			462.0	0.778	3.62	502.0	1.104	7.14	522.0	1.429	7.90	562.0	1.981	7.90	602.0	1.981	8.80
448.8			464.0	0.778	3.62	504.0	1.104	7.15	524.0	1.429	7.91	564.0	1.981	7.91	604.0	1.981	8.81
451.2			466.0	0.778	3.62	506.0	1.104	7.16	526.0	1.429	7.92	566.0	1.981	7.92	606.0	1.981	8.82
453.6			468.0	0.778	3.62	508.0	1.104	7.17	528.0	1.429	7.93	568.0	1.981	7.93	608.0	1.981	8.83
456.0			470.0	0.778	3.62	510.0	1.104	7.18	530.0	1.429	7.94	570.0	1.981	7.94	610.0	1.981	8.84
458.4			472.0	0.778	3.62	512.0	1.104	7.19	532.0	1.429	7.95	572.0	1.981	7.95	612.0	1.981	8.85
460.8			474.0	0.778	3.62	514.0	1.104	7.20	534.0	1.429	7.96	574.0	1.981	7.96	614.0	1.981	8.86
463.2			476.0	0.778	3.62	516.0	1.104	7.21	536.0	1.429	7.97	576.0	1.981	7.97	616.0	1.981	8.87
465.6			478.0	0.778	3.62	518.0	1.104	7.22	538.0	1.429	7.98	578.0	1.981	7.98	618.0	1.981	8.88
468.0			480.0	0.778	3.62	520.0	1.104	7.23	540.0	1.429	7.99</						

Table VII. (continued)

λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}
nm	$\text{W/m}^2 \cdot \text{nm}$	%	nm	$\text{W/m}^2 \cdot \text{nm}$	%	nm	$\text{W/m}^2 \cdot \text{nm}$	%	nm	$\text{W/m}^2 \cdot \text{nm}$	%	nm	$\text{W/m}^2 \cdot \text{nm}$	%	nm	$\text{W/m}^2 \cdot \text{nm}$	%
420.0	1.725	10.82	430.0	1.080	12.08	440.0	1.735	13.34	450.0	2.023	14.79	460.0	1.966	16.30	470.0	2.002	17.80
420.1	1.690	10.83	430.1	1.127	12.09	440.1	1.714	13.35	450.1	2.074	14.80	460.1	1.936	16.31	470.1	1.959	17.81
420.2	1.656	10.84	430.2	1.173	12.10	440.2	1.693	13.36	450.2	2.125	14.81	460.2	1.907	16.32	470.2	1.930	17.82
420.3	1.687	10.85	430.3	1.219	12.11	440.3	1.675	13.37	450.3	2.176	14.82	460.3	1.878	16.33	470.3	1.902	17.83
420.4	1.793	10.87	430.4	1.266	12.12	440.4	1.770	13.39	450.4	2.244	14.83	460.4	1.849	16.34	470.4	1.874	17.84
420.5	1.899	10.88	430.5	1.252	12.13	440.5	1.865	13.40	450.5	2.345	14.85	460.5	1.820	16.35	470.5	1.846	17.85
420.6	1.969	10.89	430.6	1.156	12.14	440.6	1.958	13.42	450.6	2.286	14.88	460.6	1.804	16.39	470.6	1.821	17.86
420.7	1.998	10.91	430.7	1.049	12.14	440.7	1.982	13.43	450.7	2.251	14.90	460.7	1.782	16.40	470.7	1.796	17.87
420.8	2.027	10.92	430.8	0.989	12.15	440.8	2.005	13.45	450.8	2.219	14.91	460.8	1.760	16.42	470.8	1.773	17.88
420.9	2.039	10.94	430.9	1.079	12.16	440.9	2.029	13.46	450.9	2.253	14.93	460.9	1.748	16.43	470.9	1.751	17.89
421.0	2.030	10.95	431.0	1.169	12.17	441.0	2.099	13.47	451.0	2.305	14.95	461.0	1.734	16.45	471.0	1.734	17.90
421.1	2.020	10.97	431.1	1.273	12.18	441.1	2.168	13.49	451.1	2.254	14.96	461.1	1.717	16.46	471.1	1.717	17.91
421.2	2.017	10.98	431.2	1.424	12.19	441.2	2.237	13.51	451.2	2.174	14.98	461.2	1.701	16.48	471.2	1.701	17.92
421.3	2.022	10.99	431.3	1.574	12.20	441.3	1.969	13.52	451.3	2.132	14.99	461.3	1.685	16.50	471.3	1.685	17.93
421.4	2.027	11.01	431.4	1.732	12.21	441.4	1.701	13.53	451.4	2.102	15.01	461.4	1.669	16.51	471.4	1.669	17.94
421.5	2.005	11.02	431.5	1.919	12.22	441.5	1.849	13.55	451.5	2.074	15.02	461.5	1.653	16.53	471.5	1.653	17.95
421.6	1.751	11.04	431.6	2.106	12.24	441.6	1.981	13.56	451.6	2.163	15.04	461.6	1.637	16.54	471.6	1.637	17.96
421.7	1.897	11.05	431.7	2.234	12.25	441.7	2.078	13.58	451.7	2.165	15.06	461.7	1.621	16.56	471.7	1.621	17.97
421.8	1.899	11.06	431.8	2.143	12.27	441.8	2.176	13.59	451.8	2.165	15.07	461.8	1.605	16.57	471.8	1.605	17.98
421.9	1.975	11.08	431.9	2.052	12.28	441.9	2.138	13.61	451.9	2.195	15.09	461.9	1.589	16.59	471.9	1.589	17.99
422.0	2.051	11.09	432.0	1.938	12.30	442.0	2.099	13.62	452.0	2.237	15.10	462.0	1.573	16.60	472.0	1.573	18.00
422.1	2.050	11.11	432.1	1.733	12.31	442.1	1.979	13.64	452.1	2.159	15.12	462.1	1.557	16.62	472.1	1.557	18.01
422.2	1.946	11.12	432.2	1.533	12.32	442.2	1.858	13.65	452.2	2.044	15.13	462.2	1.542	16.63	472.2	1.542	18.02
422.3	1.842	11.14	432.3	1.384	12.33	442.3	1.994	13.66	452.3	2.036	15.15	462.3	1.526	16.65	472.3	1.526	18.03
422.4	1.670	11.15	432.4	1.445	12.34	442.4	2.135	13.68	452.4	2.064	15.16	462.4	1.510	16.66	472.4	1.510	18.04
422.5	1.405	11.16	432.5	1.506	12.35	442.5	2.102	13.69	452.5	1.974	15.18	462.5	1.494	16.68	472.5	1.494	18.05
422.6	1.141	11.17	432.6	1.584	12.36	442.6	2.084	13.71	452.6	1.845	15.19	462.6	1.478	16.69	472.6	1.478	18.06
422.7	1.117	11.18	432.7	1.919	12.38	442.7	2.065	13.72	452.7	1.845	15.20	462.7	1.462	16.71	472.7	1.462	18.07
422.8	1.430	11.19	432.8	1.889	12.39	442.8	2.068	13.74	452.8	1.811	15.22	462.8	1.446	16.73	472.8	1.446	18.08
422.9	1.743	11.20	432.9	2.015	12.40	442.9	1.987	13.75	452.9	1.821	15.23	462.9	1.430	16.74	472.9	1.430	18.09
423.0	1.911	11.21	433.0	2.025	12.42	443.0	1.902	13.77	453.0	1.835	15.24	463.0	1.414	16.76	473.0	1.414	18.10
423.1	1.869	11.22	433.1	2.036	12.43	443.1	1.960	13.78	453.1	1.936	15.26	463.1	1.398	16.77	473.1	1.398	18.11
423.2	1.827	11.24	433.2	2.044	12.45	443.2	2.023	13.79	453.2	2.009	15.27	463.2	1.382	16.79	473.2	1.382	18.12
423.3	1.758	11.25	433.3	2.043	12.46	443.3	2.074	13.81	453.3	2.002	15.29	463.3	1.366	16.80	473.3	1.366	18.13
423.4	1.652	11.26	433.4	2.043	12.48	443.4	1.715	13.82	453.4	1.863	15.30	463.4	1.350	16.82	473.4	1.350	18.14
423.5	1.545	11.27	433.5	2.007	12.49	443.5	1.859	13.83	453.5	1.966	15.31	463.5	1.334	16.83	473.5	1.334	18.15
423.6	1.525	11.29	433.6	1.805	12.50	443.6	2.018	13.85	453.6	1.958	15.33	463.6	1.318	16.85	473.6	1.318	18.16
423.7	1.639	11.30	433.7	1.603	12.52	443.7	2.057	13.86	453.7	1.977	15.34	463.7	1.302	16.86	473.7	1.302	18.17
423.8	1.752	11.31	433.8	1.438	12.53	443.8	2.090	13.88	453.8	2.130	15.36	463.8	1.286	16.88	473.8	1.286	18.18
423.9	1.834	11.32	433.9	1.458	12.54	443.9	2.073	13.89	453.9	2.085	15.37	463.9	1.270	16.89	473.9	1.270	18.19
424.0	1.866	11.33	434.0	1.477	12.55	444.0	2.054	13.91	454.0	2.042	15.39	464.0	1.254	16.91	474.0	1.254	18.20
424.1	1.898	11.35	434.1	1.511	12.56	444.1	1.937	13.92	454.1	2.094	15.40	464.1	1.238	16.92	474.1	1.238	18.21
424.2	1.949	11.36	434.2	1.621	12.57	444.2	1.815	13.94	454.2	2.185	15.42	464.2	1.222	16.94	474.2	1.222	18.22
424.3	2.032	11.38	434.3	1.730	12.58	444.3	2.073	13.95	454.3	2.176	15.44	464.3	1.206	16.96	474.3	1.206	18.23
424.4	2.115	11.39	434.4	1.829	12.59	444.4	2.042	13.96	454.4	2.127	15.45	464.4	1.190	16.97	474.4	1.190	18.24
424.5	2.107	11.41	434.5	1.876	12.61	444.5	2.011	13.98	454.5	2.000	15.47	464.5	1.174	16.99	474.5	1.174	18.25
424.6	1.951	11.42	434.6	1.922	12.62	444.6	2.047	13.99	454.6	2.081	15.48	464.6	1.158	17.00	474.6	1.158	18.26
424.7	1.796	11.43	434.7	1.948	12.64	444.7	2.068	14.01	454.7	2.024	15.50	464.7	1.142	17.01	474.7	1.142	18.27
424.8	1.686	11.45	434.8	1.856	12.65	444.8	2.096	14.02	454.8	1.951	15.51	464.8	1.126	17.03	474.8	1.126	18.28
424.9	1.655	11.46	434.9	1.764	12.66	444.9	2.124	14.04	454.9	1.988	15.52	464.9	1.110	17.04	474.9	1.110	18.29
425.0	1.624	11.47	435.0	1.694	12.67	445.0	2.159	14.05	455.0	2.075	15.54	465.0	1.094	17.06	475.0	1.094	18.30
425.1	1.610	11.48	435.1	1.759	12.69	445.1	2.196	14.07	455.1	2.099	15.55	465.1	1.078	17.07	475.1	1.078	18.31
425.2	1.628	11.49	435.2	1.825	12.70	445.2	1.988	14.08	455.2	2.095	15.57	465.2	1.062	17.09	475.2	1.062	18.32
425.3	1.645	11.51	435.3	1.887	12.71	445.3	1.758	14.10	455.3	2.040	15.58	465.3	1.046	17.10	475.3	1.046	18.33
425.4	1.718	11.52	435.4	1.929	12.73	445.4	1.765	14.11	455.4	1.961	15.60	465.4	1.030	17.11	475.4	1.030	18.34
425.5	1.892	11.53	435.5	1.976	12.75	445.5	1.976	14.12	455.5	1.985	15.61	465.5	1.014	17.13	475.5	1.014	18.35
425.6	2.066	11.54	435.6	1.998	12.76	445.6	1.816	14.14	455.6	2.056	15.63	465.6	0.998	17.14	475.6	0.998	18.36
425.7	2.079	11.56	435.7	1.929	12.77	445.7	1.835	14.15	455.7	2.183	15.64	465.7	0.982	17.15	475.7	0.982	18.37
425.8	1.793	11.57	435.8	1.861	12.78	445.8	1.866	14.16	455.8	2.011	15.66	465.8	0.966	17.17	475.8	0.966	18.38
425.9	1.506	11.59	435.9	1.810	12.80	445.9	1.898	14.18	455.9	2.141	15.67	465.9	0.950	17.19	475.9	0.950	18.39
426.0	1.402	11.60	436.0	1.884	12.81	446.0	1.981	14.19	456.0	2.121	15.69	466.0	0.934	17.21	476.0	0.934	18.40
426.1	1.650	11.61	436.1	1.959	12.82	446.1	2.068	14.21	456.1	2.109	15.70	466.1	0.918	17.22	476.1	0.918	18.41
426.2	1.897	11.62	436.2	2.030	12.84	446.2	2.156	14.22	456.2	2.000	15.72	466.2	0.902	17.24	476.2	0.902	18.42
426.3	2.055	11.63	436.3	2.067	12.85	446.3	2.105	14.24	456.3	2.155	15.73	466.3	0.886	17.25	476.3	0.886	18.43
426.4	2.035	11.66	436.4	2.105	12.87	446.4	2.036	14.25	456.4	2.061	15.75	466.4	0.870	17.27	476.4</		

continued

Table VII. (continued)

λ , nm	M_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	M_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	M_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	M_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	M_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	M_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	M_{λ} , W/m ² ·nm	P_{λ} , %
440.0	2.155	20.74	500.0	1.927	22.18	510.0	1.927	23.60	520.0	1.994	24.94	530.0	1.993	26.31	540.0	1.852	27.69	550.0	1.849	29.05
440.1	2.152	20.75	500.1	1.910	22.20	510.1	1.902	23.61	520.1	1.994	24.95	530.1	1.993	26.32	540.1	1.851	27.71	550.1	1.849	29.06
440.2	2.070	20.77	500.2	1.890	22.21	510.2	1.903	23.63	520.2	2.030	24.97	530.2	2.034	26.34	540.2	1.850	27.72	550.2	1.849	29.07
440.3	1.998	20.78	500.3	1.893	22.23	510.3	1.922	23.64	520.3	2.054	25.00	530.3	2.051	26.35	540.3	1.785	27.73	550.3	1.829	29.09
440.4	2.051	20.80	500.4	1.894	22.24	510.4	2.040	23.65	520.4	2.060	25.01	530.4	2.094	26.36	540.4	1.719	27.75	550.4	1.840	29.10
440.5	2.105	20.81	500.5	1.895	22.25	510.5	2.060	23.67	520.5	1.848	25.01	530.5	2.094	26.38	540.5	1.731	27.76	550.5	1.858	29.11
440.6	2.150	20.83	500.6	1.895	22.27	510.6	2.012	23.68	520.6	1.775	25.02	530.6	2.120	26.39	540.6	1.743	27.77	550.6	1.889	29.12
440.7	2.194	20.84	500.7	1.896	22.28	510.7	1.964	23.70	520.7	1.860	25.04	530.7	2.097	26.41	540.7	1.819	27.78	550.7	1.889	29.13
440.8	2.108	20.85	500.8	1.892	22.29	510.8	1.915	23.71	520.8	1.946	25.05	530.8	2.075	26.42	540.8	1.895	27.80	550.8	1.921	29.15
440.9	2.023	20.87	500.9	2.088	22.31	510.9	1.867	23.73	520.9	1.924	25.06	530.9	2.076	26.44	540.9	1.912	27.81	550.9	1.905	29.17
491.0	1.931	20.89	501.0	2.037	22.32	511.0	1.973	23.74	521.0	1.902	25.08	531.0	2.076	26.45	541.0	1.928	27.82	551.0	1.888	29.18
491.1	1.840	20.90	501.1	1.985	22.34	511.1	2.078	23.75	521.1	1.978	25.09	531.1	2.052	26.47	541.1	1.907	27.84	551.1	1.888	29.19
491.2	1.983	20.92	501.2	1.933	22.35	511.2	2.069	23.77	521.2	2.054	25.11	531.2	2.027	26.48	541.2	1.886	27.85	551.2	1.888	29.21
491.3	2.126	20.93	501.3	1.882	22.37	511.3	2.059	23.78	521.3	2.044	25.12	531.3	1.993	26.50	541.3	1.832	27.87	551.3	1.903	29.22
491.4	2.127	20.95	501.4	1.830	22.38	511.4	2.086	23.80	521.4	2.021	25.14	531.4	1.959	26.51	541.4	1.778	27.88	551.4	1.919	29.24
491.5	2.127	20.96	501.5	1.857	22.39	511.5	2.112	23.81	521.5	1.912	25.15	531.5	1.927	26.53	541.5	1.815	27.89	551.5	1.917	29.25
491.6	2.128	20.98	501.6	1.886	22.41	511.6	2.086	23.83	521.6	1.894	25.16	531.6	1.896	26.54	541.6	1.851	27.90	551.6	1.916	29.26
491.7	1.955	20.99	501.7	1.894	22.42	511.7	2.059	23.84	521.7	1.812	25.18	531.7	1.947	26.55	541.7	1.830	27.92	551.7	1.881	29.28
491.8	1.782	21.00	501.8	1.904	22.43	511.8	2.033	23.86	521.8	1.821	25.19	531.8	1.909	26.57	541.8	1.809	27.93	551.8	1.846	29.29
491.9	1.675	21.02	501.9	1.913	22.45	511.9	2.006	23.87	521.9	1.916	25.20	531.9	2.020	26.58	541.9	1.870	27.94	551.9	1.863	29.30
492.0	1.567	21.03	502.0	1.922	22.46	512.0	1.980	23.89	522.0	2.011	25.22	532.0	2.042	26.60	542.0	1.871	27.96	552.0	1.880	29.32
492.1	1.747	21.04	502.1	1.931	22.47	512.1	1.954	23.90	522.1	1.978	25.23	532.1	1.952	26.61	542.1	1.896	27.97	552.1	1.878	29.33
492.2	1.926	21.05	502.2	1.940	22.49	512.2	1.927	23.92	522.2	1.945	25.24	532.2	1.863	26.63	542.2	1.920	27.98	552.2	1.877	29.34
492.3	1.931	21.07	502.3	1.979	22.50	512.3	1.903	23.93	522.3	1.926	25.26	532.3	1.798	26.64	542.3	1.930	28.00	552.3	1.856	29.36
492.4	1.935	21.08	502.4	2.018	22.52	512.4	1.874	23.94	522.4	1.979	25.27	532.4	1.742	26.65	542.4	1.939	28.01	552.4	1.836	29.37
492.5	2.023	21.10	502.5	1.967	22.53	512.5	1.848	23.96	522.5	1.874	25.29	532.5	1.754	26.66	542.5	1.966	28.03	552.5	1.768	29.38
492.6	2.110	21.11	502.6	1.916	22.54	512.6	1.866	23.97	522.6	1.769	25.30	532.6	1.767	26.68	542.6	1.994	28.04	552.6	1.700	29.40
492.7	2.064	21.12	502.7	1.955	22.56	512.7	1.883	23.98	522.7	1.759	25.31	532.7	1.773	26.69	542.7	1.919	28.05	552.7	1.704	29.41
492.8	2.018	21.14	502.8	1.995	22.57	512.8	1.901	24.00	522.8	1.748	25.32	532.8	1.779	26.70	542.8	1.845	28.07	552.8	1.708	29.42
492.9	2.035	21.15	502.9	2.012	22.59	512.9	1.919	24.01	522.9	1.856	25.34	532.9	1.918	26.72	542.9	1.850	28.08	552.9	1.758	29.43
493.0	2.051	21.17	503.0	2.030	22.60	513.0	1.953	24.02	523.0	1.963	25.35	533.0	2.056	26.73	543.0	1.854	28.09	553.0	1.808	29.45
493.1	2.112	21.18	503.1	2.048	22.62	513.1	1.987	24.04	523.1	1.914	25.37	533.1	2.034	26.74	543.1	1.856	28.11	553.1	1.834	29.46
493.2	2.173	21.20	503.2	2.075	22.63	513.2	1.946	24.05	523.2	1.865	25.38	533.2	2.012	26.76	543.2	1.858	28.12	553.2	1.860	29.47
493.3	2.077	21.21	503.3	2.103	22.65	513.3	1.905	24.07	523.3	1.826	25.39	533.3	2.010	26.77	543.3	1.819	28.13	553.3	1.912	29.49
493.4	1.981	21.23	503.4	1.983	22.66	513.4	1.959	24.08	523.4	1.786	25.41	533.4	2.008	26.79	543.4	1.781	28.15	553.4	1.963	29.50
493.5	2.036	21.24	503.5	2.013	22.67	513.5	1.953	24.11	523.5	1.815	25.42	533.5	1.970	26.80	543.5	1.827	28.16	553.5	1.931	29.51
493.6	2.090	21.26	503.6	1.906	22.69	513.6	1.892	24.12	523.6	1.843	25.43	533.6	1.933	26.82	543.6	1.873	28.17	553.6	1.899	29.53
493.7	1.946	21.27	503.7	1.948	22.70	513.7	1.892	24.12	523.7	1.966	25.45	533.7	1.882	26.83	543.7	1.937	28.19	553.7	1.906	29.54
493.8	1.801	21.29	503.8	1.815	22.71	513.8	1.802	24.14	523.8	2.088	25.46	533.8	1.831	26.84	543.8	1.902	28.20	553.8	1.914	29.55
493.9	1.917	21.30	503.9	1.682	22.73	513.9	1.773	24.15	523.9	2.018	25.47	533.9	1.802	26.86	543.9	2.002	28.22	553.9	1.904	29.57
494.0	2.033	21.31	504.0	1.668	22.74	514.0	1.772	24.16	524.0	1.947	25.49	534.0	1.774	26.87	544.0	2.002	28.23	554.0	1.895	29.58
494.1	2.098	21.33	504.1	1.654	22.75	514.1	1.771	24.17	524.1	1.943	25.50	534.1	1.825	26.88	544.1	1.994	28.24	554.1	1.898	29.60
494.2	2.164	21.34	504.2	1.887	22.76	514.2	1.836	24.19	524.2	1.938	25.52	534.2	1.876	26.89	544.2	1.985	28.26	554.2	1.902	29.61
494.3	2.161	21.35	504.3	2.119	22.78	514.3	1.902	24.20	524.3	1.945	25.53	534.3	1.901	26.91	544.3	1.929	28.27	554.3	1.891	29.62
494.4	2.159	21.36	504.4	2.096	22.79	514.4	1.936	24.21	524.4	1.952	25.54	534.4	1.927	26.92	544.4	1.873	28.29	554.4	1.880	29.64
494.5	2.085	21.39	504.5	2.073	22.81	514.5	1.971	24.23	524.5	2.008	25.56	534.5	1.946	26.94	544.5	1.835	28.30	554.5	1.868	29.65
494.6	2.010	21.40	504.6	2.050	22.82	514.6	1.914	24.24	524.6	2.066	25.57	534.6	1.965	26.95	544.6	1.797	28.31	554.6	1.856	29.66
494.7	2.118	21.42	504.7	2.027	22.84	514.7	1.857	24.26	524.7	2.002	25.59	534.7	1.978	26.96	544.7	1.873	28.33	554.7	1.875	29.68
494.8	2.227	21.44	504.8	2.005	22.85	514.8	1.891	24.27	524.8	1.941	25.60	534.8	1.990	26.98	544.8	1.950	28.34	554.8	1.895	29.69
494.9	2.144	21.45	504.9	1.982	22.87	514.9	1.926	24.28	524.9	1.925	25.62	534.9	2.001	26.99	544.9	1.953	28.35	554.9	1.893	29.70
495.0	2.061	21.47	505.0	1.959	22.88	515.0	1.931	24.30	525.0	1.910	25.63	535.0	2.012	27.01	545.0	1.956	28.37	555.0	1.892	29.72
495.1	2.037	21.48	505.1	1.936	22.90	515.1	1.936	24.31	525.1	1.923	25.64	535.1	2.024	27.02	545.1	1.894	28.38	555.1	1.876	29.73
495.2	2.013	21.50	505.2	1.975	22.91	515.2	1.892	24.32	525.2	1.937	25.66	535.2	2.036	27.04	545.2	1.833	28.39	555.2	1.860	29.74
495.3	2.002	21.51	505.3	2.015	22.92	515.3	1.847	24.34	525.3	1.947	25.67	535.3	2.033	27.05	545.3	1.810	28.41	555.3	1.886	29.76
495.4	1.990	21.52	505.4	2.034	22.94	515.4	1.878	24.35	525.4	1.958	25.69	535.4	2.029	27.07	545.4	1.787	28.42	555.4	1.912	29.77
495.5	1.875	21.54	505.5	2.063	22.95	515.5	1.908	24.36	525.5	1.962	25.70	535.5	2.026							

Table VII. (continued)

λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}	λ	H_{λ}	P_{λ}
nm	W/m ² ·nm	%	nm	W/m ² ·nm	%	nm	W/m ² ·nm	%	nm	W/m ² ·nm	%	nm	W/m ² ·nm	%	nm	W/m ² ·nm	%
560.0	1.712	30.38	570.0	1.822	31.70	580.0	1.886	33.05	590.0	1.900	34.38	600.0	1.816	35.69	610.0	1.915	36.98
560.1	1.717	30.39	570.1	1.840	31.71	580.1	1.844	33.06	590.1	1.961	34.39	600.1	1.814	35.71	610.1	1.830	37.00
560.2	1.723	30.40	570.2	1.858	31.73	580.2	1.803	33.08	590.2	2.023	34.40	600.2	1.813	35.72	610.2	1.745	37.01
560.3	1.796	30.42	570.3	1.815	31.74	580.3	1.803	33.09	590.3	1.930	34.42	600.3	1.836	35.73	610.3	1.735	37.02
560.4	1.870	30.43	570.4	1.772	31.75	580.4	1.803	33.10	590.4	1.837	34.43	600.4	1.860	35.75	610.4	1.758	37.03
560.5	1.876	30.44	570.5	1.758	31.77	580.5	1.817	33.12	590.5	1.816	34.44	600.5	1.846	35.76	610.5	1.771	37.05
560.6	1.883	30.46	570.6	1.745	31.78	580.6	1.836	33.13	590.6	1.796	34.46	600.6	1.832	35.77	610.6	1.784	37.06
560.7	1.881	30.47	570.7	1.713	31.79	580.7	1.846	33.14	590.7	1.805	34.47	600.7	1.815	35.79	610.7	1.797	37.07
560.8	1.880	30.48	570.8	1.680	31.80	580.8	1.855	33.16	590.8	1.815	34.48	600.8	1.799	35.80	610.8	1.811	37.09
560.9	1.850	30.50	570.9	1.678	31.82	580.9	1.887	33.17	590.9	1.831	34.50	600.9	1.846	35.81	610.9	1.842	37.10
561.0	1.820	30.51	571.0	1.675	31.83	581.0	1.919	33.18	591.0	1.846	34.51	601.0	1.893	35.83	611.0	1.873	37.11
561.1	1.808	30.52	571.1	1.729	31.84	581.1	1.943	33.20	591.1	1.860	34.52	601.1	1.831	35.84	611.1	1.811	37.13
561.2	1.795	30.54	571.2	1.782	31.85	581.2	1.967	33.21	591.2	1.874	34.54	601.2	1.774	35.85	611.2	1.750	37.14
561.3	1.779	30.55	571.3	1.840	31.87	581.3	1.928	33.23	591.3	1.848	34.55	601.3	1.738	35.86	611.3	1.731	37.15
561.4	1.762	30.56	571.4	1.897	31.88	581.4	1.889	33.24	591.4	1.822	34.56	601.4	1.702	35.88	611.4	1.726	37.16
561.5	1.734	30.57	571.5	1.889	31.89	581.5	1.839	33.25	591.5	1.839	34.58	601.5	1.691	35.89	611.5	1.724	37.18
561.6	1.705	30.59	571.6	1.881	31.91	581.6	1.788	33.27	591.6	1.855	34.59	601.6	1.679	35.90	611.6	1.722	37.19
561.7	1.768	30.60	571.7	1.899	31.92	581.7	1.802	33.28	591.7	1.857	34.60	601.7	1.711	35.91	611.7	1.722	37.20
561.8	1.830	30.61	571.8	1.917	31.93	581.8	1.816	33.29	591.8	1.859	34.62	601.8	1.743	35.93	611.8	1.723	37.21
561.9	1.877	30.62	571.9	1.939	31.95	581.9	1.850	33.30	591.9	1.850	34.63	601.9	1.744	35.94	611.9	1.720	37.22
562.0	1.925	30.64	572.0	1.961	31.96	582.0	1.883	33.32	592.0	1.842	34.64	602.0	1.746	35.95	612.0	1.716	37.24
562.1	1.896	30.65	572.1	1.981	31.98	582.1	1.880	33.33	592.1	1.841	34.66	602.1	1.725	35.96	612.1	1.722	37.25
562.2	1.868	30.67	572.2	2.000	31.99	582.2	1.877	33.35	592.2	1.840	34.67	602.2	1.704	35.98	612.2	1.729	37.26
562.3	1.846	30.68	572.3	2.021	32.00	582.3	1.851	33.36	592.3	1.869	34.68	602.3	1.696	35.99	612.3	1.738	37.27
562.4	1.824	30.69	572.4	2.042	32.02	582.4	1.824	33.37	592.4	1.899	34.70	602.4	1.688	36.00	612.4	1.778	37.29
562.5	1.859	30.71	572.5	2.009	32.03	582.5	1.831	33.38	592.5	1.856	34.71	602.5	1.758	36.01	612.5	1.761	37.30
562.6	1.894	30.72	572.6	1.977	32.05	582.6	1.838	33.40	592.6	1.813	34.72	602.6	1.745	36.02	612.6	1.745	37.31
562.7	1.886	30.73	572.7	1.943	32.06	582.7	1.844	33.41	592.7	1.836	34.74	602.7	1.723	36.04	612.7	1.749	37.32
562.8	1.878	30.75	572.8	1.909	32.08	582.8	1.860	33.42	592.8	1.860	34.75	602.8	1.818	36.05	612.8	1.752	37.34
562.9	1.891	30.76	572.9	1.886	32.09	582.9	1.910	33.44	592.9	1.813	34.76	602.9	1.802	36.06	612.9	1.729	37.35
563.0	1.905	30.77	573.0	1.863	32.10	583.0	1.960	33.45	593.0	1.766	34.78	603.0	1.787	36.08	613.0	1.706	37.36
563.1	1.911	30.79	573.1	1.877	32.12	583.1	1.902	33.47	593.1	1.784	34.79	603.1	1.787	36.09	613.1	1.690	37.37
563.2	1.926	30.80	573.2	1.891	32.13	583.2	1.844	33.48	593.2	1.797	34.80	603.2	1.787	36.10	613.2	1.674	37.39
563.3	1.928	30.81	573.3	1.835	32.14	583.3	1.835	33.49	593.3	1.790	34.81	603.3	1.786	36.12	613.3	1.710	37.40
563.4	1.929	30.83	573.4	1.845	32.16	583.4	1.827	33.51	593.4	1.783	34.83	603.4	1.784	36.13	613.4	1.745	37.41
563.5	1.899	30.84	573.5	1.833	32.17	583.5	1.862	33.52	593.5	1.826	34.84	603.5	1.806	36.14	613.5	1.679	37.42
563.6	1.869	30.86	573.6	1.821	32.18	583.6	1.896	33.53	593.6	1.870	34.85	603.6	1.829	36.15	613.6	1.612	37.44
563.7	1.847	30.87	573.7	1.879	32.20	583.7	1.907	33.55	593.7	1.874	34.87	603.7	1.829	36.17	613.7	1.659	37.45
563.8	1.825	30.88	573.8	1.937	32.21	583.8	1.919	33.56	593.8	1.879	34.88	603.8	1.829	36.18	613.8	1.706	37.46
563.9	1.837	30.89	573.9	1.941	32.22	583.9	1.934	33.57	593.9	1.877	34.89	603.9	1.787	36.19	613.9	1.722	37.47
564.0	1.849	30.91	574.0	1.945	32.24	584.0	1.949	33.59	594.0	1.876	34.91	604.0	1.745	36.21	614.0	1.738	37.48
564.1	1.880	30.92	574.1	1.915	32.25	584.1	1.974	33.60	594.1	1.862	34.92	604.1	1.727	36.22	614.1	1.731	37.50
564.2	1.912	30.94	574.2	1.886	32.27	584.2	1.999	33.62	594.2	1.848	34.93	604.2	1.709	36.23	614.2	1.723	37.51
564.3	1.886	30.95	574.3	1.878	32.28	584.3	1.970	33.63	594.3	1.879	34.95	604.3	1.706	36.24	614.3	1.757	37.52
564.4	1.824	30.96	574.4	1.870	32.29	584.4	1.910	33.64	594.4	1.910	34.96	604.4	1.704	36.26	614.4	1.792	37.53
564.5	1.834	30.98	574.5	1.902	32.31	584.5	1.893	33.66	594.5	1.894	34.97	604.5	1.712	36.27	614.5	1.763	37.55
564.6	1.846	30.99	574.6	1.933	32.32	584.6	1.846	33.67	594.6	1.877	34.99	604.6	1.720	36.28	614.6	1.736	37.56
564.7	1.840	31.00	574.7	1.893	32.33	584.7	1.850	33.68	594.7	1.852	35.00	604.7	1.738	36.29	614.7	1.738	37.57
564.8	1.835	31.01	574.8	1.852	32.35	584.8	1.854	33.70	594.8	1.827	35.01	604.8	1.756	36.30	614.8	1.741	37.58
564.9	1.851	31.03	574.9	1.833	32.36	584.9	1.836	33.71	594.9	1.841	35.03	604.9	1.785	36.32	614.9	1.764	37.60
565.0	1.867	31.04	575.0	1.814	32.37	585.0	1.818	33.72	595.0	1.855	35.04	605.0	1.814	36.33	615.0	1.788	37.61
565.1	1.885	31.05	575.1	1.814	32.39	585.1	1.832	33.74	595.1	1.840	35.05	605.1	1.817	36.34	615.1	1.782	37.62
565.2	1.903	31.07	575.2	1.815	32.40	585.2	1.846	33.75	595.2	1.824	35.07	605.2	1.820	36.36	615.2	1.776	37.64
565.3	1.865	31.08	575.3	1.853	32.41	585.3	1.857	33.76	595.3	1.816	35.08	605.3	1.807	36.37	615.3	1.727	37.65
565.4	1.827	31.10	575.4	1.868	32.43	585.4	1.868	33.78	595.4	1.807	35.09	605.4	1.795	36.38	615.4	1.678	37.66
565.5	1.815	31.11	575.5	1.929	32.44	585.5	1.828	33.79	595.5	1.772	35.11	605.5	1.795	36.40	615.5	1.686	37.67
565.6	1.803	31.12	575.6	1.968	32.45	585.6	1.787	33.80	595.6	1.737	35.12	605.6	1.795	36.41	615.6	1.695	37.68
565.7	1.778	31.13	575.7	1.876	32.47	585.7	1.814	33.82	595.7	1.728	35.13	605.7	1.860	36.42	615.7	1.733	37.70
565.8	1.754	31.15	575.8	1.969	32.48	585.8	1.840	33.83	595.8	1.720	35.14	605.8	1.925	36.44	615.8	1.771	37.71
565.9	1.797	31.16	575.9	1.951	32.50	585.9	1.898	33.84	595.9	1.803	35.16	605.9	1.900	36.45	615.9	1.702	37.72
566.0	1.841	31.17	576.0	1.932	32.51	586.0	1.957	33.86	596.0	1.886	35.17	606.0	1.875	36.46	616.0	1.634	37.73
566.1	1.814	31.19	576.1	1.883	32.52	586.1	1.944	33.87	596.1	1.911	35.18	606.1	1.887	36.48	616.1	1.618	37.75
566.2	1.787	31.20	576.2	1.834	32.54	586.2	1.932	33.88	596.2	1.936	35.20	606.2	1.900	36.49	616.2	1.603	37.76
566.3	1.787	31.21	576.3	1.824	32.55	586.3	1.879	33.90	596.3	1.963	35.21	606.3	1.850	36.50	616.3	1.627	37.77
566.4	1.825	31.22	576.4	1.855	32.56	586.4	1.826	33.91	596.4	1							

Table VII. (concluded)

λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %	λ , nm	H_{λ} , W/m ² ·nm	P_{λ} , %
640.0	1.685	40.68	660.0	1.680	42.98	700.0	1.495	47.45	900.0	0.889	64.03	1250.0	0.461	80.30	1750.0	0.191	91.30	2250.0	0.077	95.61
640.2	1.686	40.71	660.4	1.657	43.03	702.0	1.471	47.66	902.0	0.891	64.16	1255.0	0.451	80.47	1755.0	0.193	91.37	2255.0	0.077	95.64
640.4	1.686	40.73	660.8	1.633	43.07	704.0	1.447	47.87	904.0	0.881	64.29	1260.0	0.432	80.62	1760.0	0.188	91.44	2260.0	0.072	95.66
640.6	1.570	40.75	661.2	1.613	43.12	706.0	1.420	48.08	906.0	0.891	64.42	1265.0	0.426	80.78	1765.0	0.188	91.51	2265.0	0.073	95.69
640.8	1.557	40.78	661.6	1.599	43.16	708.0	1.427	48.29	908.0	0.854	64.54	1270.0	0.427	80.93	1770.0	0.180	91.57	2270.0	0.068	95.71
641.0	1.581	40.80	662.0	1.623	43.21	710.0	1.455	48.50	910.0	0.865	64.66	1275.0	0.441	81.09	1775.0	0.176	91.64	2275.0	0.065	95.73
641.2	1.620	40.82	662.4	1.604	43.26	712.0	1.434	48.71	912.0	0.873	64.79	1280.0	0.424	81.25	1780.0	0.174	91.70	2280.0	0.070	95.76
641.4	1.642	40.84	662.8	1.593	43.30	714.0	1.421	48.91	914.0	0.865	64.91	1285.0	0.426	81.39	1785.0	0.173	91.76	2285.0	0.067	95.78
641.6	1.650	40.87	663.2	1.561	43.35	716.0	1.406	49.11	916.0	0.860	65.04	1290.0	0.431	81.55	1790.0	0.172	91.82	2290.0	0.067	95.81
641.8	1.643	40.89	663.6	1.553	43.39	718.0	1.393	49.32	918.0	0.861	65.16	1295.0	0.427	81.70	1795.0	0.176	91.88	2295.0	0.066	95.83
642.0	1.642	40.92	664.0	1.654	43.44	720.0	1.367	49.51	920.0	0.835	65.28	1300.0	0.427	81.86	1800.0	0.172	91.95	2300.0	0.067	95.86
642.2	1.670	40.94	664.4	1.570	43.49	722.0	1.396	49.71	922.0	0.828	65.40	1305.0	0.418	82.01	1805.0	0.171	92.01	2305.0	0.064	95.88
642.4	1.727	40.96	664.8	1.540	43.53	724.0	1.406	49.91	924.0	0.834	65.52	1310.0	0.409	82.16	1810.0	0.163	92.07	2310.0	0.063	95.90
642.6	1.714	40.99	665.2	1.627	43.58	726.0	1.392	50.11	926.0	0.835	65.64	1315.0	0.406	82.30	1815.0	0.163	92.13	2315.0	0.059	95.92
642.8	1.634	41.01	665.6	1.640	43.63	728.0	1.366	50.31	928.0	0.845	65.76	1320.0	0.407	82.45	1820.0	0.162	92.19	2320.0	0.054	95.94
643.0	1.609	41.04	666.0	1.636	43.67	730.0	1.347	50.51	930.0	0.840	65.88	1325.0	0.402	82.60	1825.0	0.159	92.24	2325.0	0.056	95.96
643.2	1.581	41.06	666.4	1.559	43.72	732.0	1.362	50.70	932.0	0.848	66.00	1330.0	0.396	82.74	1830.0	0.153	92.36	2335.0	0.060	96.01
643.4	1.597	41.08	666.8	1.642	43.76	734.0	1.359	50.90	934.0	0.835	66.13	1335.0	0.392	82.88	1835.0	0.159	92.30	2340.0	0.056	95.98
643.6	1.652	41.10	667.2	1.619	43.81	736.0	1.363	51.10	936.0	0.820	66.24	1340.0	0.390	83.02	1840.0	0.156	92.41	2345.0	0.059	96.03
643.8	1.699	41.13	667.6	1.628	43.86	738.0	1.332	51.29	938.0	0.826	66.36	1345.0	0.387	83.16	1845.0	0.154	92.47	2350.0	0.061	96.05
644.0	1.714	41.15	668.0	1.598	43.90	740.0	1.316	51.48	940.0	0.821	66.48	1350.0	0.380	83.30	1850.0	0.151	92.52	2350.0	0.062	96.07
644.2	1.698	41.18	668.4	1.650	43.95	742.0	1.302	51.67	942.0	0.806	66.60	1355.0	0.375	83.43	1855.0	0.148	92.58	2355.0	0.064	96.09
644.4	1.669	41.20	668.8	1.534	44.00	744.0	1.306	51.85	944.0	0.822	66.71	1360.0	0.372	83.57	1860.0	0.146	92.63	2360.0	0.066	96.12
644.6	1.597	41.23	669.2	1.625	44.04	746.0	1.308	52.04	946.0	0.817	66.83	1365.0	0.364	83.70	1865.0	0.146	92.68	2365.0	0.062	96.14
644.8	1.564	41.25	669.6	1.640	44.08	748.0	1.314	52.23	948.0	0.801	66.95	1370.0	0.363	83.83	1870.0	0.138	92.73	2370.0	0.062	96.16
645.0	1.599	41.27	670.0	1.692	44.13	750.0	1.304	52.42	950.0	0.803	67.06	1375.0	0.363	83.96	1875.0	0.138	92.78	2375.0	0.058	96.18
645.2	1.671	41.29	670.4	1.561	44.18	752.0	1.297	52.61	952.0	0.796	67.18	1380.0	0.359	84.09	1880.0	0.143	92.83	2380.0	0.056	96.20
645.4	1.660	41.32	670.8	1.536	44.22	754.0	1.294	52.80	954.0	0.785	67.29	1385.0	0.359	84.22	1885.0	0.141	92.88	2385.0	0.051	96.22
645.6	1.708	41.34	671.2	1.585	44.27	756.0	1.292	52.98	956.0	0.789	67.41	1390.0	0.355	84.35	1890.0	0.140	92.93	2390.0	0.060	96.25
645.8	1.746	41.37	671.6	1.550	44.31	758.0	1.283	53.16	958.0	0.796	67.52	1395.0	0.353	84.47	1895.0	0.141	92.98	2395.0	0.059	96.27
646.0	1.697	41.39	672.0	1.528	44.36	760.0	1.269	53.35	960.0	0.792	67.63	1400.0	0.349	84.60	1900.0	0.136	93.03	2400.0	0.055	96.29
646.2	1.616	41.42	672.4	1.520	44.40	762.0	1.272	53.53	962.0	0.791	67.75	1405.0	0.346	84.72	1905.0	0.139	93.08	2405.0	0.056	96.31
646.4	1.615	41.44	672.8	1.538	44.45	764.0	1.255	53.71	964.0	0.788	67.86	1410.0	0.343	84.85	1910.0	0.141	93.13	2410.0	0.057	96.33
646.6	1.609	41.46	673.2	1.520	44.49	766.0	1.251	53.89	966.0	0.787	67.97	1415.0	0.341	84.97	1915.0	0.142	93.18	2415.0	0.054	96.35
646.8	1.636	41.49	673.6	1.598	44.54	768.0	1.235	54.07	968.0	0.798	68.09	1420.0	0.340	85.09	1920.0	0.137	93.23	2420.0	0.058	96.37
647.0	1.656	41.51	674.0	1.602	44.58	770.0	1.226	54.25	970.0	0.783	68.20	1425.0	0.345	85.22	1925.0	0.135	93.28	2425.0	0.055	96.39
647.2	1.626	41.53	674.4	1.664	44.63	772.0	1.217	54.42	972.0	0.783	68.31	1430.0	0.335	85.34	1930.0	0.135	93.33	2430.0	0.059	96.41
647.4	1.595	41.56	674.8	1.633	44.68	774.0	1.216	54.60	974.0	0.767	68.43	1435.0	0.329	85.46	1935.0	0.134	93.38	2435.0	0.055	96.43
647.6	1.599	41.58	675.2	1.577	44.72	776.0	1.197	54.77	976.0	0.783	68.54	1440.0	0.325	85.57	1940.0	0.132	93.43	2440.0	0.052	96.45
647.8	1.621	41.60	675.6	1.605	44.77	778.0	1.218	54.94	978.0	0.781	68.65	1445.0	0.317	85.69	1945.0	0.130	93.47	2445.0	0.050	96.47
648.0	1.615	41.63	676.0	1.578	44.81	780.0	1.223	55.12	980.0	0.772	68.76	1450.0	0.322	85.81	1950.0	0.129	93.52	2450.0	0.051	96.48
648.2	1.629	41.65	676.4	1.558	44.86	782.0	1.217	55.29	982.0	0.774	68.87	1455.0	0.306	85.92	1955.0	0.124	93.57	2455.0	0.052	96.50
648.4	1.625	41.67	676.8	1.566	44.90	784.0	1.218	55.47	984.0	0.776	68.98	1460.0	0.316	86.03	1960.0	0.128	93.61	2460.0	0.049	96.52
648.6	1.591	41.69	677.2	1.604	44.95	786.0	1.213	55.64	986.0	0.758	69.09	1465.0	0.311	86.14	1965.0	0.128	93.66	2465.0	0.054	96.54
648.8	1.542	41.72	677.6	1.634	45.00	788.0	1.205	55.82	988.0	0.767	69.20	1470.0	0.311	86.25	1970.0	0.127	93.70	2470.0	0.049	96.56
649.0	1.518	41.74	678.0	1.537	45.04	790.0	1.185	55.99	990.0	0.763	69.31	1475.0	0.307	86.37	1975.0	0.132	93.75	2475.0	0.054	96.58
649.2	1.443	41.76	678.4	1.559	45.09	792.0	1.160	56.15	992.0	0.754	69.42	1480.0	0.303	86.47	1980.0	0.127	93.80	2480.0	0.051	96.59
649.4	1.420	41.78	678.8	1.564	45.13	794.0	1.177	56.32	994.0	0.756	69.53	1485.0	0.299	86.58	1985.0	0.125	93.84	2485.0	0.048	96.61
649.6	1.509	41.80	679.2	1.573	45.18	796.0	1.172	56.49	996.0	0.743	69.64	1490.0	0.304	86.69	1990.0	0.123	93.89	2490.0	0.055	96.63
649.8	1.556	41.82	679.6	1.574	45.22	798.0	1.163	56.66	998.0	0.743	69.74	1495.0	0.301	86.80	1995.0	0.125	93.93	2495.0	0.047	96.65
650.0	1.577	41.85	680.0	1.578	45.27	800.0	1.154	56.83	1000.0	0.741	69.85	1500.0	0.297	86.91	2000.0	0.118	93.97			
650.2	1.615	41.87	680.4	1.605	45.31	802.0	1.150	56.99	1005.0	0.725	70.11	1505.0	0.296	87.01	2005.0	0.112	94.01			
650.4	1.654	41.89	680.8	1.554	45.36	804.0	1.148	57.16	1010.0	0.723	70.37	1510.0	0.292	87.12	2010.0	0.116	94.06			
650.6	1.640	41.92	681.2	1.527	45.40	806.0	1.141	57.32	1015.0	0.722	70.63	1515.0	0.292	87.22	2015.0	0.114	94.10			
650.8	1.606	41.94	681.6	1.581	45.45	808.0	1.120	57.48	1020.0	0.716	70.89	1520.0	0.288	87.33	2020.0	0.115	94.14			
651.0	1.600	41.96	682.0	1.578	45.49	8														

The second procedure utilized the Cary spectroradiometer (see Fig. 6). An uncollimated high-intensity tungsten lamp was installed in the sample side 50 cm from the limiting aperture. Polarization of radiation from this lamp was less than 2%. The window was placed into and out of the sample light path and the ratio of signals noted. Measurements were performed at selected wavelengths for incidence angles of 0°, 20°, and 30°. After correction was made for the difference in optical path lengths for air and fused silica, the average disagreement in transmittance between the two sets of measurements made with different equipment was ± 0.006 . After a smooth curve was fit through each set of data, the disagreement averaged ± 0.004 . There was no measurable difference in spectral transmittance between either of the windows used during the solar measurements. The tabulation of transmittance as a function of wavelength and incidence angle is given in Table V.

B. Measurement of Mirror Reflectance

The spectral reflectance of the mirror was determined by a multiple reflection technique to obtain the highest degree of accuracy. The method used is somewhat similar to the techniques of Strong²⁸ or Bennett and Koehler,²⁹ which enable measurements to be made to a precision of $\pm 0.1\%$.

During solar measurements, the incidence angle on the turning mirror ranged from 25° to 55°, depending on the solar zenith angle. The mirror reflectance exhibits at most a 2% variation in reflectance for unpolarized radiation over this range. Because of this small variation with incidence angle, measurements were performed only at the angle most commonly encountered during the measurements (35°), and Fresnel's equations for an aluminum-air interface were used to obtain the variation with angle.

The apparatus used in the reflectance measurements is shown in Fig. 16. Essentially the same instrumentation was used as for the window transmittance measurements described previously. Two mirrors, prepared under identical conditions, were placed opposite each other at a separation distance of 6 cm. Collimated radiation from a xenon-arc source was directed horizontally into the integrating sphere to obtain the reference signal level, V_R , and then directed onto the lower mirror at an angle of incidence of 35°. At this angle of incidence, the light beam undergoes seven reflections between the two mirrors before entering the integrating sphere, which was rotated downward to accept the beam, as shown in Fig. 16. The intensity of the radiation entering the integrating sphere is indicated by the sample signal level, V_s . The reflectance R of the mirror is obtained by the seventh root of the signal ratio:

$$R = (V_s/V_R)^{1/7}. \quad (A1)$$

Measurements were performed at each wavelength with the beam polarized both parallel (R_p) and normal (R_n) to the plane of incidence through the use of a Nicol prism. The spectral reflectance ρ_λ for natural or

unpolarized radiation was calculated from the arithmetic mean of the two component reflectances

$$\rho_\lambda = (R_p + R_n)/2. \quad (A2)$$

The results of the measurements are shown in Fig. 17 for 0° incidence and are listed in Table VI for incidence angles 0°, 35°, 45°, and 55°.

In preparing the mirror, the aluminum was deposited at a rate of 50 nm/sec at a pressure of 5×10^{-6} Torr. The total thickness was approximately 500 nm. Also shown in Fig. 17 are the data of Hass³⁰ for a fresh aluminum surface evaporated at an extremely high evaporation rate (100 nm/sec) in good vacuum (1×10^{-5} Torr) and the data of Bennett *et al.*³¹ for an aluminum surface deposited at a slower rate and approximately the same pressure level. In general, the present measurements are very close to the previously published data. The 2% decrease in the uv near 300 nm is felt to be due to the lower rate of deposition and higher pressure during evaporation than used by Hass. The data of Hass and Bennett were used beyond 1200 nm since the low output of the xenon-arc source at longer wavelengths resulted in a low signal level. The mirror reflectance indicated by the solid curve in Fig. 17 and the listing in Table VI are felt to be within $\pm 0.5\%$ of the true reflectance.

Appendix B. Extraterrestrial Spectral Irradiance from 300 nm to 2500 nm

The data presented in this section are representative of the degree of resolution attained in the experimental results over the 300–2500 nm spectral range. All values are on an absolute spectral irradiance scale based on the Standard of Spectral Irradiance and can be used to calculate total or average irradiance within arbitrary spectral bands.

A view of the Fraunhofer absorption in the solar atmosphere is shown in Fig. 18. The spectral region from 300 nm to 700 nm has considerably greater structure than the 700–2500 nm region and is shown in higher detail. The identities of selected major absorption lines are indicated in the spectra.

A tabular listing of the data is given in Table VII. The values of extraterrestrial spectral irradiance, $H_{\lambda 0}$, are averaged over the printout interval. A listing of the percentage P_λ of the total solar irradiance at wavelengths shorter than the respective wavelength is also shown.

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November 13, 1969

EXTRATERRESTIAL SOLAR SPECTRAL IRRADIANCES

TABLE OF CONTENTS

- I. Correspondence from Sender
- II. Program for Listing Tape
- III. Output of Program

From Ames - C
To Goddard
Code 601
National

Subject:

Reference:

1. The
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Moffett Field, California 94035
October 23, 1969

From Jones - Code STP
To Goddard Space Flight Center
Code 601
National Space Science Data Center

Subject: Data on extraterrestrial solar spectral irradiance

Reference: Arvesen, John C.; Griffin, Roy M., Jr.; and Pearson, E.
Douglas, Jr.: "Determination of Extraterrestrial Solar
Spectral Irradiance from a Research Aircraft," Applied
Optics, 8, 11, 1969

1. The magnetic tape, which is being sent under separate cover, contains the extraterrestrial solar spectral irradiance listing reported in the referenced paper (copy enclosed). The computer output of a program which reads the tape and prints the listing is also enclosed. The program is written in Fortran IV and was run on the IBM 7040/7094 DCS.

2. If you require any further information, please feel free to contact us.

John C. Arvesen

John C. Arvesen
Aerospace Engineer

Encs.

1. Referenced paper(1)
2. Computer output (1)
3. Magnetic tape (under separate cover)

	0001	REC	1. LENGTH	84
	0002	REC	2. LENGTH	84
	0003	REC	3. LENGTH	84
SOLAR SPECTRAL IRRADIANCE	0004	REC	4. LENGTH	84
AT	0005	REC	5. LENGTH	84
N SOLAR DISTANCE	0006	REC	6. LENGTH	84
- 2500NM)	0007	REC	7. LENGTH	84
*****	0008	REC	8. LENGTH	84
	0009	REC	9. LENGTH	84
	0010	REC	10. LENGTH	84
C. GRIFFIN, ROY N. JR. AND	0011	REC	11. LENGTH	84
--DETERMINATION OF EXTRA--	0012	REC	12. LENGTH	84
IRRAL IRRADIANCE FROM A RESEARCH	0013	REC	13. LENGTH	84
PTICS, VOL 8, NO 11, 1969.	0014	REC	14. LENGTH	84
	0015	REC	15. LENGTH	84
TRIAL PERCENTAGE OF TOTAL	0016	REC	16. LENGTH	84
DIANCE, IRRADIANCE AT	0017	REC	17. LENGTH	84
ANOMETER SHORTER WAVELENGTH	0018	REC	18. LENGTH	84
*****	0019	REC	19. LENGTH	84
1.15	0020	REC	20. LENGTH	84
1.16	0021	REC	21. LENGTH	84
1.18	0022	REC	22. LENGTH	84
1.19	0023	REC	23. LENGTH	84
1.21	0024	REC	24. LENGTH	84
1.23	0025	REC	25. LENGTH	84
1.25	0026	REC	26. LENGTH	84
1.26	0027	REC	27. LENGTH	84
1.28	0028	REC	28. LENGTH	84
1.30	0029	REC	29. LENGTH	84
1.31	0030	REC	30. LENGTH	84
1.33	0031	REC	31. LENGTH	84
1.35	0032	REC	32. LENGTH	84

CORR55I
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305.2	0.628	1.36	0033
305.6	0.646	1.38	0034
306.0	0.490	1.40	0035
306.4	0.557	1.41	0036
306.8	0.622	1.43	0037
307.2	0.718	1.45	0038
307.6	0.746	1.47	0039
308.0	0.672	1.49	0040
308.4	0.631	1.51	0041
308.8	0.620	1.53	0042
309.2	0.533	1.54	0043
309.6	0.514	1.56	0044
310.0	0.536	1.57	0045
310.4	0.540	1.59	0046
310.8	0.707	1.61	0047
311.2	0.792	1.63	0048
311.6	0.793	1.65	0049
312.0	0.706	1.67	0050
312.4	0.659	1.69	0051
312.8	0.747	1.71	0052
313.2	0.759	1.73	0053
313.6	0.706	1.75	0054
314.0	0.755	1.78	0055
314.4	0.774	1.80	0056
314.8	0.851	1.82	0057
315.2	0.822	1.84	0058
315.6	0.800	1.86	0059
316.0	0.629	1.88	0060
316.4	0.576	1.90	0061
316.8	0.744	1.92	0062
317.2	0.841	1.94	0063
317.6	0.917	1.97	0064

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1.38	0034	REC	34. LENGTH	84
1.40	0035	REC	35. LENGTH	84
1.41	0036	REC	36. LENGTH	84
1.43	0037	REC	37. LENGTH	84
1.45	0038	REC	38. LENGTH	84
1.47	0039	REC	39. LENGTH	84
1.49	0040	REC	40. LENGTH	84
1.51	0041	REC	41. LENGTH	84
1.53	0042	REC	42. LENGTH	84
1.54	0043	REC	43. LENGTH	84
1.56	0044	REC	44. LENGTH	84
1.57	0045	REC	45. LENGTH	84
1.59	0046	REC	46. LENGTH	84
1.61	0047	REC	47. LENGTH	84
1.63	0048	REC	48. LENGTH	84
1.65	0049	REC	49. LENGTH	84
1.67	0050	REC	50. LENGTH	84
1.69	0051	REC	51. LENGTH	84
1.71	0052	REC	52. LENGTH	84
1.73	0053	REC	53. LENGTH	84
1.75	0054	REC	54. LENGTH	84
1.78	0055	REC	55. LENGTH	84
1.80	0056	REC	56. LENGTH	84
1.82	0057	REC	57. LENGTH	84
1.84	0058	REC	58. LENGTH	84
1.86	0059	REC	59. LENGTH	84
1.88	0060	REC	60. LENGTH	84
1.90	0061	REC	61. LENGTH	84
1.92	0062	REC	62. LENGTH	84
1.94	0063	REC	63. LENGTH	84
1.97	0064	REC	64. LENGTH	84

96.29	3450	REC	1. LENGTH	84
96.31	3451	REC	2. LENGTH	84
96.33	3452	REC	3. LENGTH	84
96.35	3453	REC	4. LENGTH	84
96.37	3454	REC	5. LENGTH	84
96.39	3455	REC	6. LENGTH	84
96.41	3456	REC	7. LENGTH	84
96.43	3457	REC	8. LENGTH	84
96.45	3458	REC	9. LENGTH	84
96.47	3459	REC	10. LENGTH	84
96.48	3460	REC	11. LENGTH	84
96.50	3461	REC	12. LENGTH	84
96.52	3462	REC	13. LENGTH	84
96.54	3463	REC	14. LENGTH	84
96.56	3464	REC	15. LENGTH	84
96.58	3465	REC	16. LENGTH	84
96.59	3466	REC	17. LENGTH	84
96.61	3467	REC	18. LENGTH	84
96.63	3468	REC	19. LENGTH	84
96.65	3469	REC	20. LENGTH	84

2400.0	0.055	96.29	3450
2405.0	0.056	96.31	3451
2410.0	0.057	96.33	3452
2415.0	0.054	96.35	3453
2420.0	0.056	96.37	3454
2425.0	0.055	96.39	3455
2430.0	0.059	96.41	3456
2435.0	0.055	96.43	3457
2440.0	0.052	96.45	3458
2445.0	0.050	96.47	3459
2450.0	0.051	96.48	3460
2455.0	0.052	96.50	3461
2460.0	0.049	96.52	3462
2465.0	0.054	96.54	3463
2470.0	0.049	96.56	3464
2475.0	0.054	96.58	3465
2480.0	0.051	96.59	3466
2485.0	0.048	96.61	3467
2490.0	0.055	96.63	3468
2495.0	0.047	96.65	3469

EXTRATERRESTRIAL SOLAR SPECTRAL IRRADIANCE

AT

EARTH'S MEAN SOLAR DISTANCE

(300 - 2500NM)

REFERENCE***** ARVESEN, JOHN C, GRIFFIN, ROY N, JR, AND
PEARSON, B DOUGLAS, JR, --DETERMINATION OF EXTRA-
TERRESTRIAL SOLAR SPECTRAL IRRADIANCE FROM A RESEARCH
AIRCRAFT--, APPLIED OPTICS, VOL 8, NO 11, 1969.

WAVELENGTH, NANOMETER	EXTRATERRESTRIAL SPECTRAL IRRADIANCE, WATT/SQMETER-NANOMETER	PERCENTAGE OF TOTAL IRRADIANCE AT SHORTER WAVELENGTH
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300.0	0.545	1.13
300.4	0.504	1.16
300.8	0.546	1.16
301.2	0.616	1.19
301.6	0.749	1.21
302.0	0.576	1.23
302.4	0.485	1.25
302.8	0.541	1.26
303.2	0.590	1.28
303.6	0.635	1.30
304.0	0.563	1.31
304.4	0.522	1.33
304.8	0.552	1.35
305.2	0.628	1.36
305.6	0.646	1.38
306.0	0.490	1.40
306.4	0.557	1.41
306.8	0.622	1.43
307.2	0.718	1.45
307.6	0.746	1.47
308.0	0.672	1.49
308.4	0.631	1.51
308.8	0.620	1.53
309.2	0.533	1.54
309.6	0.514	1.56
310.0	0.536	1.57
310.4	0.540	1.59
310.8	0.707	1.61
311.2	0.792	1.63
311.6	0.793	1.65
312.0	0.706	1.67
312.4	0.659	1.69
312.8	0.747	1.71
313.2	0.759	1.73
313.6	0.706	1.75
314.0	0.755	1.79
314.4	0.774	1.80
314.8	0.651	1.82
315.2	0.822	1.84

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59	315.6	0.800	1.86	0059
60	316.0	0.629	1.88	0060
61	316.4	0.576	1.90	0061
62	316.8	0.744	1.92	0062
63	317.2	0.841	1.94	0063
64	317.6	0.917	1.97	0064
65	318.0	0.805	1.99	0065
66	318.4	0.620	2.01	0066
67	318.8	0.735	2.03	0067
68	319.2	0.887	2.06	0068
69	319.6	0.721	2.08	0069
70	320.0	0.786	2.10	0070
71	320.4	0.861	2.12	0071
72	320.8	0.925	2.15	0072
73	321.2	0.867	2.18	0073
74	321.6	0.691	2.20	0074
75	322.0	0.775	2.22	0075
76	322.4	0.856	2.24	0076
77	322.8	0.747	2.27	0077
78	323.2	0.619	2.29	0078
79	323.6	0.657	2.30	0079
80	324.0	0.620	2.32	0080
81	324.4	0.687	2.34	0081
82	324.8	0.697	2.36	0082
83	325.2	0.668	2.38	0083
84	325.6	0.773	2.40	0084
85	326.0	1.029	2.43	0085
86	326.4	1.189	2.46	0086
87	326.8	1.155	2.49	0087
88	327.2	1.183	2.53	0088
89	327.6	1.034	2.56	0089
90	328.0	1.004	2.58	0090
91	328.4	0.928	2.61	0091
92	328.8	1.000	2.64	0092
93	329.2	1.009	2.67	0093
94	329.6	1.193	2.70	0094
95	330.0	1.223	2.74	0095
96	330.4	1.187	2.77	0096
97	330.8	1.014	2.80	0097
98	331.2	0.993	2.83	0098
99	331.6	1.022	2.86	0099
100	332.0	1.057	2.89	0100
101	332.4	1.074	2.92	0101
102	332.8	1.030	2.95	0102
103	333.2	1.062	2.98	0103
104	333.6	1.076	3.01	0104
105	334.0	0.930	3.04	0105
106	334.4	1.047	3.07	0106
107	334.8	1.179	3.10	0107
108	335.2	1.033	3.13	0108
109	335.6	1.119	3.16	0109
110	336.0	1.131	3.20	0110
111	336.4	0.698	3.22	0111
112	336.8	0.909	3.24	0112
113	337.2	0.864	3.27	0113
114	337.6	0.827	3.30	0114
115	338.0	0.988	3.32	0115
116	338.4	0.938	3.35	0116

117	338.8	1.039	3.38	0117
118	339.2	1.091	3.41	0118
119	339.6	1.005	3.44	0119
120	340.0	1.089	3.47	0120
121	340.2	1.185	3.48	0121
122	340.4	1.175	3.50	0122
123	340.6	1.147	3.52	0123
124	340.8	1.078	3.53	0124
125	341.0	1.013	3.55	0125
126	341.2	1.028	3.56	0126
127	341.4	1.084	3.58	0127
128	341.6	1.072	3.59	0128
129	341.8	0.899	3.61	0129
130	342.0	0.972	3.62	0130
131	342.2	1.104	3.64	0131
132	342.4	1.207	3.65	0132
133	342.6	1.133	3.67	0133
134	342.8	1.047	3.69	0134
135	343.0	1.035	3.70	0135
136	343.2	1.101	3.72	0136
137	343.4	1.207	3.73	0137
138	343.6	1.163	3.75	0138
139	343.8	1.095	3.77	0139
140	344.0	1.023	3.78	0140
141	344.2	0.905	3.80	0141
142	344.4	0.670	3.81	0142
143	344.6	0.624	3.82	0143
144	344.8	0.770	3.83	0144
145	345.0	0.899	3.84	0145
146	345.2	1.073	3.85	0146
147	345.4	1.189	3.87	0147
148	345.6	1.068	3.88	0148
149	345.8	1.095	3.90	0149
150	346.0	1.085	3.92	0150
151	346.2	0.897	3.93	0151
152	346.4	0.818	3.94	0152
153	346.6	0.901	3.95	0153
154	346.8	1.065	3.97	0154
155	347.0	0.988	3.98	0155
156	347.2	1.099	4.00	0156
157	347.4	1.163	4.01	0157
158	347.6	1.082	4.03	0158
159	347.8	0.873	4.05	0159
160	348.0	0.815	4.06	0160
161	348.2	0.909	4.07	0161
162	348.4	1.046	4.08	0162
163	348.6	0.979	4.10	0163
164	348.8	0.915	4.11	0164
165	349.0	1.020	4.13	0165
166	349.2	1.020	4.14	0166
167	349.4	0.900	4.15	0167
168	349.6	0.870	4.17	0168
169	349.8	0.942	4.18	0169
170	350.0	1.015	4.19	0170
171	350.2	1.088	4.21	0171
172	350.4	1.160	4.22	0172
173	350.6	1.210	4.24	0173
174	350.8	1.173	4.26	0174

175	351.0	1.059	4.28	0175
176	351.2	1.058	4.29	0176
177	351.4	1.008	4.31	0177
178	351.6	0.899	4.32	0178
179	351.8	1.002	4.33	0179
180	352.0	1.128	4.35	0180
181	352.2	1.096	4.36	0181
182	352.4	0.936	4.38	0182
183	352.6	0.821	4.39	0183
184	352.8	0.935	4.40	0184
185	353.0	1.046	4.42	0185
186	353.2	1.108	4.43	0186
187	353.4	1.113	4.45	0187
188	353.6	1.132	4.47	0188
189	353.8	1.165	4.48	0189
190	354.0	1.196	4.50	0190
191	354.2	1.203	4.52	0191
192	354.4	1.197	4.53	0192
193	354.6	1.192	4.55	0193
194	354.8	1.195	4.57	0194
195	355.0	1.213	4.58	0195
196	355.2	1.166	4.60	0196
197	355.4	1.052	4.62	0197
198	355.6	1.046	4.63	0198
199	355.8	1.051	4.65	0199
200	356.0	1.038	4.66	0200
201	356.2	1.055	4.68	0201
202	356.4	1.097	4.69	0202
203	356.6	0.861	4.71	0203
204	356.8	0.768	4.72	0204
205	357.0	0.780	4.73	0205
206	357.2	0.861	4.74	0206
207	357.4	0.939	4.75	0207
208	357.6	0.987	4.77	0208
209	357.8	0.943	4.78	0209
210	358.0	0.726	4.80	0210
211	358.2	0.575	4.80	0211
212	358.4	0.563	4.81	0212
213	358.6	0.582	4.82	0213
214	358.8	0.755	4.83	0214
215	359.0	0.976	4.84	0215
216	359.2	1.117	4.86	0216
217	359.4	1.165	4.87	0217
218	359.6	1.210	4.89	0218
219	359.8	1.207	4.91	0219
220	360.0	1.160	4.93	0220
221	360.2	1.109	4.94	0221
222	360.4	1.044	4.96	0222
223	360.6	0.979	4.97	0223
224	360.8	0.960	4.99	0224
225	361.0	0.971	5.00	0225
226	361.2	0.981	5.01	0226
227	361.4	0.941	5.03	0227
228	361.6	0.799	5.04	0228
229	361.8	0.675	5.06	0229
230	362.0	0.778	5.06	0230
231	362.2	0.968	5.07	0231
232	362.4	1.103	5.09	0232

233	362.6	1.166	5.10	0233
234	362.8	1.206	5.12	0234
235	363.0	1.123	5.14	0235
236	363.2	1.013	5.16	0236
237	363.4	1.012	5.17	0237
238	363.6	1.124	5.18	0238
239	363.8	1.212	5.20	0239
240	364.0	1.152	5.22	0240
241	364.2	1.060	5.23	0241
242	364.4	1.003	5.25	0242
243	364.6	0.986	5.26	0243
244	364.8	0.986	5.28	0244
245	365.0	1.094	5.29	0245
246	365.2	1.231	5.31	0246
247	365.4	1.315	5.33	0247
248	365.6	1.351	5.35	0248
249	365.8	1.397	5.37	0249
250	366.0	1.410	5.39	0250
251	366.2	1.373	5.41	0251
252	366.4	1.333	5.42	0252
253	366.6	1.300	5.44	0253
254	366.8	1.265	5.46	0254
255	367.0	1.259	5.48	0255
256	367.2	1.284	5.50	0256
257	367.4	1.305	5.52	0257
258	367.6	1.248	5.54	0258
259	367.8	1.154	5.55	0259
260	368.0	1.096	5.57	0260
261	368.2	1.116	5.58	0261
262	368.4	1.147	5.60	0262
263	368.6	1.218	5.62	0263
264	368.8	1.312	5.64	0264
265	369.0	1.366	5.66	0265
266	369.2	1.316	5.67	0266
267	369.4	1.258	5.69	0267
268	369.6	1.279	5.71	0268
269	369.8	1.355	5.73	0269
270	370.0	1.371	5.75	0270
271	370.2	1.194	5.77	0271
272	370.4	0.996	5.78	0272
273	370.6	0.973	5.80	0273
274	370.8	1.091	5.81	0274
275	371.0	1.209	5.83	0275
276	371.2	1.327	5.85	0276
277	371.4	1.443	5.87	0277
278	371.6	1.322	5.89	0278
279	371.8	0.981	5.90	0279
280	372.0	0.721	5.92	0280
281	372.2	0.951	5.93	0281
282	372.4	1.220	5.94	0282
283	372.6	1.228	5.96	0283
284	372.8	1.176	5.98	0284
285	373.0	1.061	5.99	0285
286	373.2	0.931	6.01	0286
287	373.4	0.781	6.02	0287
288	373.6	0.677	6.03	0288
289	373.8	0.908	6.04	0289
290	374.0	1.175	6.06	0290

291	374.2	1.093	6.07	0291
292	374.4	0.926	6.08	0292
293	374.6	0.735	6.10	0293
294	374.8	0.623	6.11	0294
295	375.0	0.714	6.12	0295
296	375.2	1.051	6.13	0296
297	375.4	1.219	6.15	0297
298	375.6	1.109	6.17	0298
299	375.8	1.059	6.18	0299
300	376.0	1.137	6.20	0300
301	376.2	1.191	6.21	0301
302	376.4	1.190	6.23	0302
303	376.6	1.196	6.25	0303
304	376.8	1.221	6.27	0304
305	377.0	1.263	6.28	0305
306	377.2	1.332	6.30	0306
307	377.4	1.390	6.32	0307
308	377.6	1.418	6.34	0308
309	377.8	1.455	6.36	0309
310	378.0	1.515	6.38	0310
311	378.2	1.547	6.41	0311
312	378.4	1.513	6.43	0312
313	378.6	1.426	6.45	0313
314	378.8	1.228	6.47	0314
315	379.0	1.073	6.48	0315
316	379.2	1.062	6.50	0316
317	379.4	1.056	6.51	0317
318	379.6	1.019	6.53	0318
319	379.8	1.011	6.54	0319
320	380.0	1.108	6.56	0320
321	380.1	1.164	6.57	0321
322	380.2	1.216	6.58	0322
323	380.3	1.260	6.58	0323
324	380.4	1.304	6.59	0324
325	380.5	1.348	6.60	0325
326	380.6	1.387	6.61	0326
327	380.7	1.406	6.62	0327
328	380.8	1.426	6.63	0328
329	380.9	1.446	6.64	0329
330	381.0	1.446	6.65	0330
331	381.1	1.341	6.66	0331
332	381.2	1.237	6.67	0332
333	381.3	1.132	6.68	0333
334	381.4	1.043	6.69	0334
335	381.5	1.057	6.70	0335
336	381.6	1.071	6.71	0336
337	381.7	1.085	6.71	0337
338	381.8	1.092	6.72	0338
339	381.9	1.017	6.73	0339
340	382.0	0.942	6.74	0340
341	382.1	0.867	6.74	0341
342	382.2	0.794	6.75	0342
343	382.3	0.744	6.75	0343
344	382.4	0.695	6.76	0344
345	382.5	0.645	6.76	0345
346	382.6	0.596	6.77	0346
347	382.7	0.640	6.77	0347
348	382.8	0.684	6.78	0348

1	952.0	0.796	67.18	3146
2	954.0	0.785	67.29	3147
3	956.0	0.799	67.41	3148
4	958.0	0.796	67.52	3149
5	960.0	0.792	67.63	3150
6	962.0	0.791	67.75	3151
7	964.0	0.788	67.86	3152
8	966.0	0.780	67.97	3153
9	968.0	0.798	68.09	3154
10	970.0	0.783	68.20	3155
11	972.0	0.783	68.31	3156
12	974.0	0.767	68.43	3157
13	976.0	0.783	68.54	3158
14	978.0	0.781	68.65	3159
15	980.0	0.772	68.76	3160
16	982.0	0.774	68.87	3161
17	984.0	0.776	68.98	3162
18	986.0	0.758	69.09	3163
19	988.0	0.767	69.20	3164
20	990.0	0.763	69.31	3165
21	992.0	0.754	69.42	3166
22	994.0	0.756	69.53	3167
23	996.0	0.743	69.64	3168
24	998.0	0.743	69.74	3169
25	1000.0	0.741	69.85	3170
26	1005.0	0.725	70.11	3171
27	1010.0	0.723	70.37	3172
28	1015.0	0.722	70.63	3173
29	1020.0	0.716	70.89	3174
30	1025.0	0.725	71.15	3175
31	1030.0	0.700	71.41	3176
32	1035.0	0.695	71.65	3177
33	1040.0	0.679	71.90	3178
34	1045.0	0.686	72.15	3179
35	1050.0	0.668	72.39	3180
36	1055.0	0.659	72.63	3181
37	1060.0	0.650	72.86	3182
38	1065.0	0.648	73.10	3183
39	1070.0	0.640	73.33	3184
40	1075.0	0.631	73.55	3185
41	1080.0	0.621	73.78	3186
42	1085.0	0.614	74.00	3187
43	1090.0	0.612	74.22	3188
44	1095.0	0.599	74.44	3189
45	1100.0	0.608	74.66	3190
46	1105.0	0.601	74.87	3191
47	1110.0	0.602	75.09	3192
48	1115.0	0.588	75.30	3193
49	1120.0	0.578	75.51	3194
50	1125.0	0.568	75.72	3195
51	1130.0	0.565	75.92	3196
52	1135.0	0.562	76.12	3197
53	1140.0	0.556	76.32	3198
54	1145.0	0.554	76.52	3199
55	1150.0	0.543	76.73	3200
56	1155.0	0.552	76.93	3201
57	1160.0	0.538	77.11	3202
58	1165.0	0.528	77.30	3203

59	1170.0	0.531	77.50	3204
60	1175.0	0.523	77.68	3205
61	1180.0	0.512	77.87	3206
62	1185.0	0.509	78.05	3207
63	1190.0	0.508	78.23	3208
64	1195.0	0.498	78.42	3209
65	1200.0	0.492	78.59	3210
66	1205.0	0.489	78.77	3211
67	1210.0	0.484	78.95	3212
68	1215.0	0.494	79.12	3213
69	1220.0	0.475	79.29	3214
70	1225.0	0.474	79.47	3215
71	1230.0	0.476	79.64	3216
72	1235.0	0.468	79.80	3217
73	1240.0	0.466	79.97	3218
74	1245.0	0.453	80.14	3219
75	1250.0	0.461	80.30	3220
76	1255.0	0.451	80.47	3221
77	1260.0	0.432	80.62	3222
78	1265.0	0.426	80.78	3223
79	1270.0	0.427	80.93	3224
80	1275.0	0.441	81.09	3225
81	1280.0	0.424	81.25	3226
82	1285.0	0.426	81.39	3227
83	1290.0	0.431	81.55	3228
84	1295.0	0.427	81.70	3229
85	1300.0	0.427	81.86	3230
86	1305.0	0.418	82.01	3231
87	1310.0	0.409	82.16	3232
88	1315.0	0.406	82.30	3233
89	1320.0	0.407	82.45	3234
90	1325.0	0.402	82.60	3235
91	1330.0	0.396	82.74	3236
92	1335.0	0.392	82.88	3237
93	1340.0	0.390	83.02	3238
94	1345.0	0.387	83.16	3239
95	1350.0	0.380	83.30	3240
96	1355.0	0.375	83.43	3241
97	1360.0	0.372	83.57	3242
98	1365.0	0.364	83.70	3243
99	1370.0	0.363	83.83	3244
100	1375.0	0.363	83.96	3245
101	1380.0	0.359	84.09	3246
102	1385.0	0.359	84.22	3247
103	1390.0	0.354	84.35	3248
104	1395.0	0.353	84.47	3249
105	1400.0	0.349	84.60	3250
106	1405.0	0.346	84.72	3251
107	1410.0	0.343	84.85	3252
108	1415.0	0.341	84.97	3253
109	1420.0	0.340	85.09	3254
110	1425.0	0.345	85.22	3255
111	1430.0	0.335	85.34	3256
112	1435.0	0.329	85.46	3257
113	1440.0	0.325	85.57	3258
114	1445.0	0.317	85.69	3259
115	1450.0	0.322	85.81	3260
116	1455.0	0.306	85.92	3261

117	1460.0	0.316	86.03	3262
118	1465.0	0.311	86.14	3263
119	1470.0	0.311	86.25	3264
120	1475.0	0.307	86.37	3265
121	1480.0	0.303	86.47	3266
122	1485.0	0.299	86.58	3267
123	1490.0	0.304	86.69	3268
124	1495.0	0.301	86.80	3269
125	1500.0	0.297	86.91	3270
126	1505.0	0.296	87.01	3271
127	1510.0	0.292	87.12	3272
128	1515.0	0.292	87.22	3273
129	1520.0	0.288	87.33	3274
130	1525.0	0.292	87.43	3275
131	1530.0	0.284	87.54	3276
132	1535.0	0.277	87.64	3277
133	1540.0	0.278	87.73	3278
134	1545.0	0.277	87.84	3279
135	1550.0	0.276	87.93	3280
136	1555.0	0.275	88.03	3281
137	1560.0	0.273	88.13	3282
138	1565.0	0.267	88.23	3283
139	1570.0	0.264	88.32	3284
140	1575.0	0.263	88.42	3285
141	1580.0	0.259	88.51	3286
142	1585.0	0.256	88.61	3287
143	1590.0	0.250	88.70	3288
144	1595.0	0.251	88.79	3289
145	1600.0	0.252	88.88	3290
146	1605.0	0.247	88.97	3291
147	1610.0	0.249	89.05	3292
148	1615.0	0.248	89.14	3293
149	1620.0	0.245	89.23	3294
150	1625.0	0.249	89.32	3295
151	1630.0	0.246	89.41	3296
152	1635.0	0.242	89.50	3297
153	1640.0	0.238	89.58	3298
154	1645.0	0.239	89.67	3299
155	1650.0	0.238	89.75	3300
156	1655.0	0.238	89.84	3301
157	1660.0	0.237	89.93	3302
158	1665.0	0.233	90.01	3303
159	1670.0	0.232	90.09	3304
160	1675.0	0.224	90.18	3305
161	1680.0	0.225	90.26	3306
162	1685.0	0.223	90.34	3307
163	1690.0	0.223	90.42	3308
164	1695.0	0.218	90.50	3309
165	1700.0	0.221	90.57	3310
166	1705.0	0.216	90.65	3311
167	1710.0	0.207	90.73	3312
168	1715.0	0.216	90.80	3313
169	1720.0	0.209	90.88	3314
170	1725.0	0.200	90.95	3315
171	1730.0	0.194	91.03	3316
172	1735.0	0.193	91.09	3317
173	1740.0	0.195	91.16	3318
174	1745.0	0.189	91.23	3319

175	1750.0	0.191	91.30	3320
176	1755.0	0.193	91.37	3321
177	1760.0	0.188	91.44	3322
178	1765.0	0.186	91.51	3323
179	1770.0	0.180	91.57	3324
180	1775.0	0.176	91.64	3325
181	1780.0	0.174	91.70	3326
182	1785.0	0.173	91.76	3327
183	1790.0	0.172	91.82	3328
184	1795.0	0.176	91.88	3329
185	1800.0	0.172	91.95	3330
186	1805.0	0.171	92.01	3331
187	1810.0	0.163	92.07	3332
188	1815.0	0.163	92.13	3333
189	1820.0	0.162	92.19	3334
190	1825.0	0.159	92.24	3335
191	1830.0	0.159	92.30	3336
192	1835.0	0.153	92.36	3337
193	1840.0	0.156	92.41	3338
194	1845.0	0.154	92.47	3339
195	1850.0	0.151	92.52	3340
196	1855.0	0.148	92.58	3341
197	1860.0	0.146	92.63	3342
198	1865.0	0.146	92.68	3343
199	1870.0	0.138	92.73	3344
200	1875.0	0.138	92.78	3345
201	1880.0	0.143	92.83	3346
202	1885.0	0.141	92.88	3347
203	1890.0	0.140	92.93	3348
204	1895.0	0.141	92.98	3349
205	1900.0	0.136	93.03	3350
206	1905.0	0.139	93.08	3351
207	1910.0	0.141	93.13	3352
208	1915.0	0.139	93.18	3353
209	1920.0	0.137	93.23	3354
210	1925.0	0.135	93.28	3355
211	1930.0	0.135	93.33	3356
212	1935.0	0.134	93.38	3357
213	1940.0	0.132	93.43	3358
214	1945.0	0.130	93.47	3359
215	1950.0	0.129	93.52	3360
216	1955.0	0.124	93.57	3361
217	1960.0	0.128	93.61	3362
218	1965.0	0.128	93.66	3363
219	1970.0	0.127	93.70	3364
220	1975.0	0.132	93.75	3365
221	1980.0	0.127	93.80	3366
222	1985.0	0.125	93.84	3367
223	1990.0	0.123	93.89	3368
224	1995.0	0.125	93.93	3369
225	2000.0	0.118	93.97	3370
226	2005.0	0.112	94.01	3371
227	2010.0	0.116	94.06	3372
228	2015.0	0.114	94.10	3373
229	2020.0	0.115	94.14	3374
230	2025.0	0.115	94.18	3375
231	2030.0	0.112	94.22	3376
232	2035.0	0.109	94.26	3377

233	2040.0	0.109	94.30	3378
234	2045.0	0.110	94.34	3379
235	2050.0	0.106	94.38	3380
236	2055.0	0.104	94.41	3381
237	2060.0	0.102	94.45	3382
238	2065.0	0.097	94.49	3383
239	2070.0	0.101	94.52	3384
240	2075.0	0.103	94.56	3385
241	2080.0	0.099	94.60	3386
242	2085.0	0.098	94.63	3387
243	2090.0	0.100	94.67	3388
244	2095.0	0.093	94.70	3389
245	2100.0	0.095	94.74	3390
246	2105.0	0.095	94.77	3391
247	2110.0	0.089	94.80	3392
248	2115.0	0.092	94.83	3393
249	2120.0	0.089	94.87	3394
250	2125.0	0.089	94.90	3395
251	2130.0	0.089	94.93	3396
252	2135.0	0.087	94.96	3397
253	2140.0	0.092	95.00	3398
254	2145.0	0.084	95.03	3399
255	2150.0	0.083	95.06	3400
256	2155.0	0.081	95.09	3401
257	2160.0	0.084	95.12	3402
258	2165.0	0.082	95.15	3403
259	2170.0	0.084	95.18	3404
260	2175.0	0.079	95.21	3405
261	2180.0	0.077	95.23	3406
262	2185.0	0.069	95.26	3407
263	2190.0	0.074	95.28	3408
264	2195.0	0.075	95.31	3409
265	2200.0	0.084	95.34	3410
266	2205.0	0.077	95.37	3411
267	2210.0	0.077	95.40	3412
268	2215.0	0.074	95.42	3413
269	2220.0	0.079	95.45	3414
270	2225.0	0.077	95.48	3415
271	2230.0	0.077	95.51	3416
272	2235.0	0.066	95.53	3417
273	2240.0	0.061	95.55	3418
274	2245.0	0.073	95.58	3419
275	2250.0	0.077	95.61	3420
276	2255.0	0.077	95.64	3421
277	2260.0	0.072	95.65	3422
278	2265.0	0.073	95.69	3423
279	2270.0	0.066	95.71	3424
280	2275.0	0.065	95.73	3425
281	2280.0	0.070	95.76	3426
282	2285.0	0.067	95.78	3427
283	2290.0	0.067	95.81	3428
284	2295.0	0.066	95.83	3429
285	2300.0	0.067	95.86	3430
286	2305.0	0.064	95.88	3431
287	2310.0	0.063	95.90	3432
288	2315.0	0.059	95.92	3433
289	2320.0	0.054	95.94	3434
290	2325.0	0.056	95.96	3435

291	2330.0	0.056	96.98	3436
292	2335.0	0.060	96.01	3437
293	2340.0	0.059	96.03	3438
294	2345.0	0.061	96.05	3439
295	2350.0	0.062	96.07	3440
296	2355.0	0.064	96.09	3441
297	2360.0	0.066	96.12	3442
298	2365.0	0.062	96.14	3443
299	2370.0	0.062	96.16	3444
300	2375.0	0.058	96.18	3445
301	2380.0	0.056	96.20	3446
302	2385.0	0.061	96.22	3447
303	2390.0	0.060	96.25	3448
304	2395.0	0.059	96.27	3449
305	2400.0	0.055	96.29	3450
306	2405.0	0.056	96.31	3451
307	2410.0	0.057	96.33	3452
308	2415.0	0.054	96.35	3453
309	2420.0	0.058	96.37	3454
310	2425.0	0.055	96.39	3455
311	2430.0	0.059	96.41	3456
312	2435.0	0.055	96.43	3457
313	2440.0	0.052	96.45	3458
314	2445.0	0.050	96.47	3459
315	2450.0	0.051	96.48	3460
316	2455.0	0.052	96.50	3461
317	2460.0	0.049	96.52	3462
318	2465.0	0.054	96.54	3463
319	2470.0	0.049	96.56	3464
320	2475.0	0.054	96.58	3465
321	2480.0	0.051	96.59	3466
322	2485.0	0.048	96.61	3467
323	2490.0	0.055	96.63	3468
324	2495.0	0.047	96.65	3469
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